

Scattering amplitudes and their classical limits

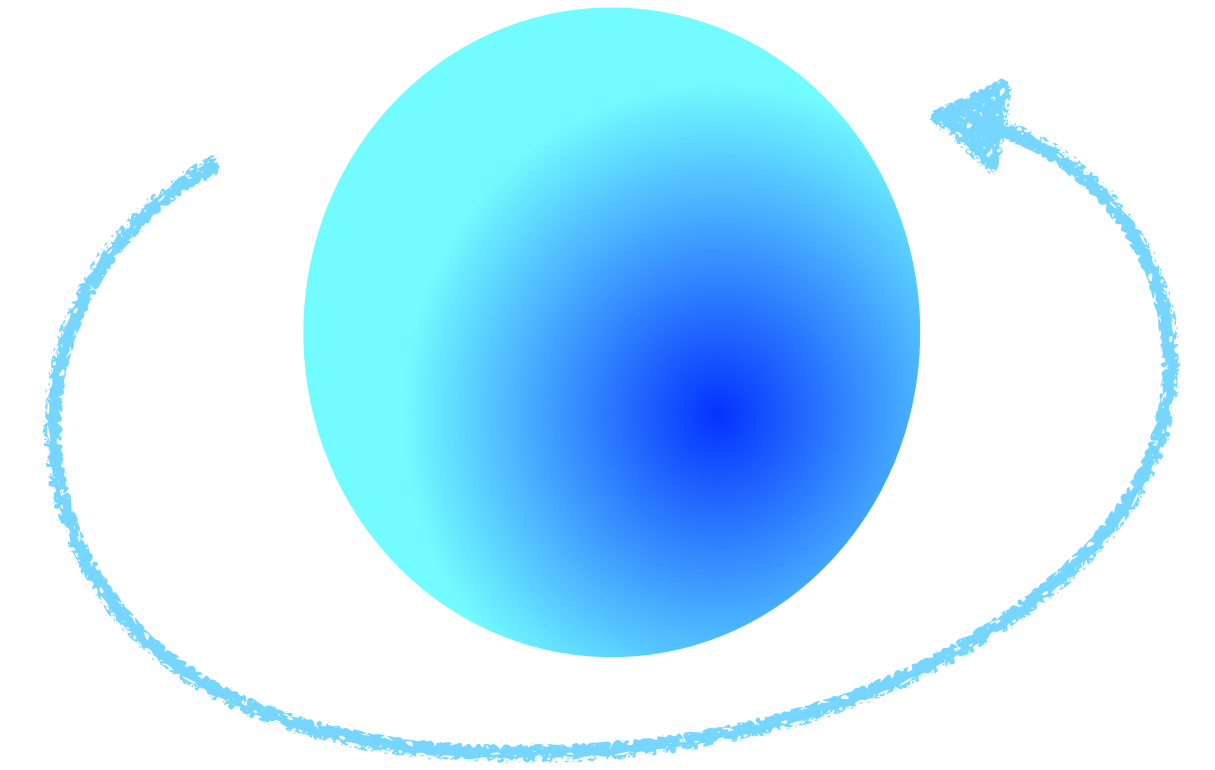
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PART I

Spin-dependent



Scattering amplitudes
and their classical limits

Spinor-helicity variables

$$p_{a\dot{a}} = p_\mu (\sigma^\mu)_{a\dot{a}}$$

$$\det(p) = 0$$

Massless : $p_{a\dot{a}} = -\lambda_a \bar{\lambda}_{\dot{a}} = -|p]_a \langle p|_{\dot{a}}$

Massive : $p_{a\dot{a}} = -\lambda_{aI} \bar{\lambda}^I_{\dot{a}} = -|p]_{aI} \langle p|^I_{\dot{a}}$

$$\det(p) = m^2$$

$$\det(\lambda) = m$$

Spinor-helicity variables

"Should I learn how to use them?"

QFT course, 1st semester

$$\Psi(x) = \int \widetilde{d^3p} \left[e^{ipx} u_I(p) b^I + e^{-ipx} v^I(p) b_I^\dagger \right]$$

$$\{b^I(p), b_J^\dagger(q)\} = (2\pi)^3 (2\omega_{\vec{p}}) \delta^3(\vec{p} - \vec{q}) \delta^I_J$$



Wigner's "little group" SU(2)
for massive particles

QFT course, 1st semester

[Peskin-Schoeder, chapter 3]

There are two linearly independent solutions for $u(p)$,

$$u^s(p) = \begin{pmatrix} \sqrt{p \cdot \sigma} \xi^s \\ \sqrt{p \cdot \bar{\sigma}} \xi^s \end{pmatrix}, \quad s = 1, 2$$

which we normalize according to

$$\bar{u}^r(p) u^s(p) = 2m \delta^{rs} \quad \text{or} \quad u^{r\dagger}(p) u^s(p) = 2E_{\mathbf{p}} \delta^{rs}.$$

δ^{rs} makes no sense!

$\delta^{\bar{r}s}$ or δ^r_s will do.

QFT course, 1st semester

Little-group covariant notation:

$$p_{a\dot{a}} = -\lambda_{aI}\bar{\lambda}^I_{\dot{a}}$$

$$u_I(p) = \begin{pmatrix} \lambda_{aI} \\ \bar{\lambda}^{\dot{a}}_I \end{pmatrix}, \quad v^I(p) = \begin{pmatrix} -\lambda_{aJ} \\ \bar{\lambda}^{\dot{a}}_J \end{pmatrix} \epsilon^{JI},$$
$$\bar{u}^I(p) = \begin{pmatrix} \lambda^{Ia} & \bar{\lambda}^I_{\dot{a}} \end{pmatrix}, \quad \bar{v}_I(p) = \epsilon_{IJ} \begin{pmatrix} \lambda^{Ja} & -\bar{\lambda}^J_{\dot{a}} \end{pmatrix}.$$

Fine. But, are they useful in practical computations?

Helicity basis

[Jacob-Wick 1959]

3d "helicity basis"

$$\hat{k} = (\sin \theta \cos \phi, \sin \theta \sin \phi, \cos \theta)$$

$$\hat{k} \cdot \vec{\sigma} = |k_+) (k_+| - |k_-) (k_-|,$$

$$|k_+) = \begin{pmatrix} c_{\theta/2} \\ e^{+i\phi} s_{\theta/2} \end{pmatrix}, \quad |k_-) = \begin{pmatrix} -e^{-i\phi} s_{\theta/2} \\ c_{\theta/2} \end{pmatrix}.$$

Helicity basis

(1+3)d "helicity basis"

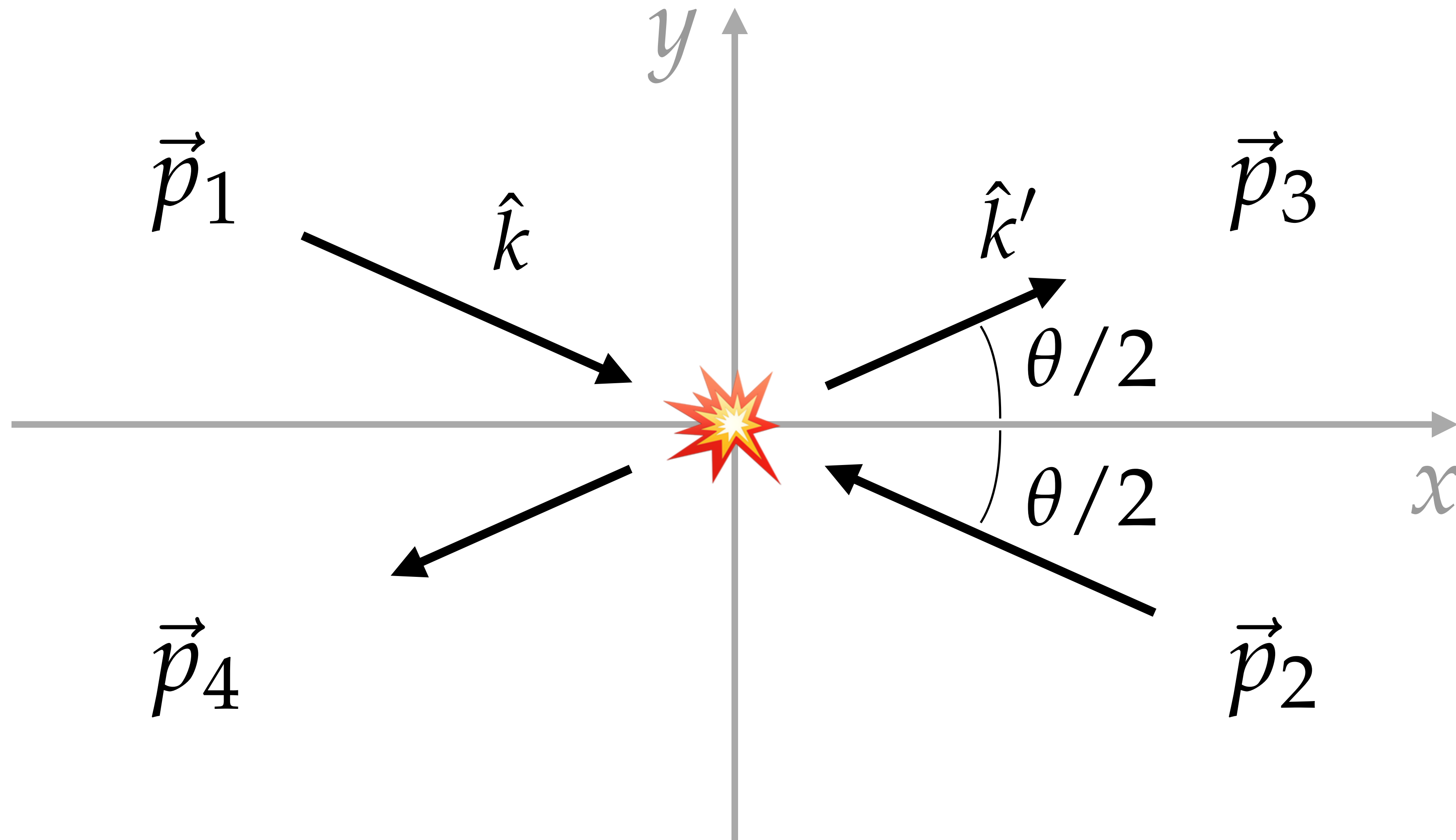
$$E = m \cosh \rho, \quad \vec{k} = m \sinh \rho \hat{k}.$$

$$\lambda_{aI} = \sqrt{m} \begin{pmatrix} e^{-\rho/2} |k_+ \rangle & e^{+\rho/2} |k_- \rangle \end{pmatrix}, \quad \bar{\lambda}^I_{\dot{a}} = \sqrt{m} \begin{pmatrix} e^{-\rho/2} \langle k_+ | \\ e^{+\rho/2} \langle k_- | \end{pmatrix},$$

$$\bar{\lambda}^{\dot{a}}_I = \sqrt{m} \begin{pmatrix} e^{+\rho/2} |k_+ \rangle & e^{-\rho/2} |k_- \rangle \end{pmatrix}, \quad \lambda^{Ia} = \sqrt{m} \begin{pmatrix} e^{+\rho/2} \langle k_+ | \\ e^{-\rho/2} \langle k_- | \end{pmatrix}.$$

A QED process

$$e^- e^+ \rightarrow \mu^+ \mu^-$$



A QED process

$$e^- e^+ \rightarrow \mu^+ \mu^-$$

$$\mathcal{M}^{I_3 I_4}_{I_1 I_2} = \frac{e^2}{s} \left[\bar{u}^{I_3}(p_3) \gamma_\mu v^{I_4}(p_4) \right] \left[\bar{v}_{I_2}(p_2) \gamma^\mu u_{I_1}(p_1) \right]$$

↑
(4x4) matrix

$$\mathcal{M} = -\frac{2e^2}{s} \left(\langle \mathbf{31} \rangle [\mathbf{42}] + \langle \mathbf{42} \rangle [\mathbf{31}] + \langle \mathbf{41} \rangle [\mathbf{32}] + \langle \mathbf{32} \rangle [\mathbf{41}] \right)$$

↑ ↑
(2x2) blocks

A QED process

$$e^- e^+ \rightarrow \mu^+ \mu^-$$

$$\begin{aligned} \widetilde{\mathcal{M}} &= -\frac{1}{2m_e m_\mu} (\langle \mathbf{31} \rangle [\mathbf{42}] + \langle \mathbf{42} \rangle [\mathbf{31}] + \langle \mathbf{41} \rangle [\mathbf{32}] + \langle \mathbf{32} \rangle [\mathbf{41}]) \\ &= \begin{pmatrix} c_\theta & -ich_e s_\theta & -ich_e s_\theta & c_\theta \\ -ich_\mu s_\theta & ch_\mu ch_e (c_\theta + 1) & ch_\mu ch_e (c_\theta - 1) & -ich_\mu s_\theta \\ -ich_\mu s_\theta & ch_\mu ch_e (c_\theta - 1) & ch_\mu ch_e (c_\theta + 1) & -ich_\mu s_\theta \\ c_\theta & -ich_e s_\theta & -ich_e s_\theta & c_\theta \end{pmatrix} \end{aligned}$$

$$m_e \cosh \rho_e = E = m_\mu \cosh \rho_\mu$$

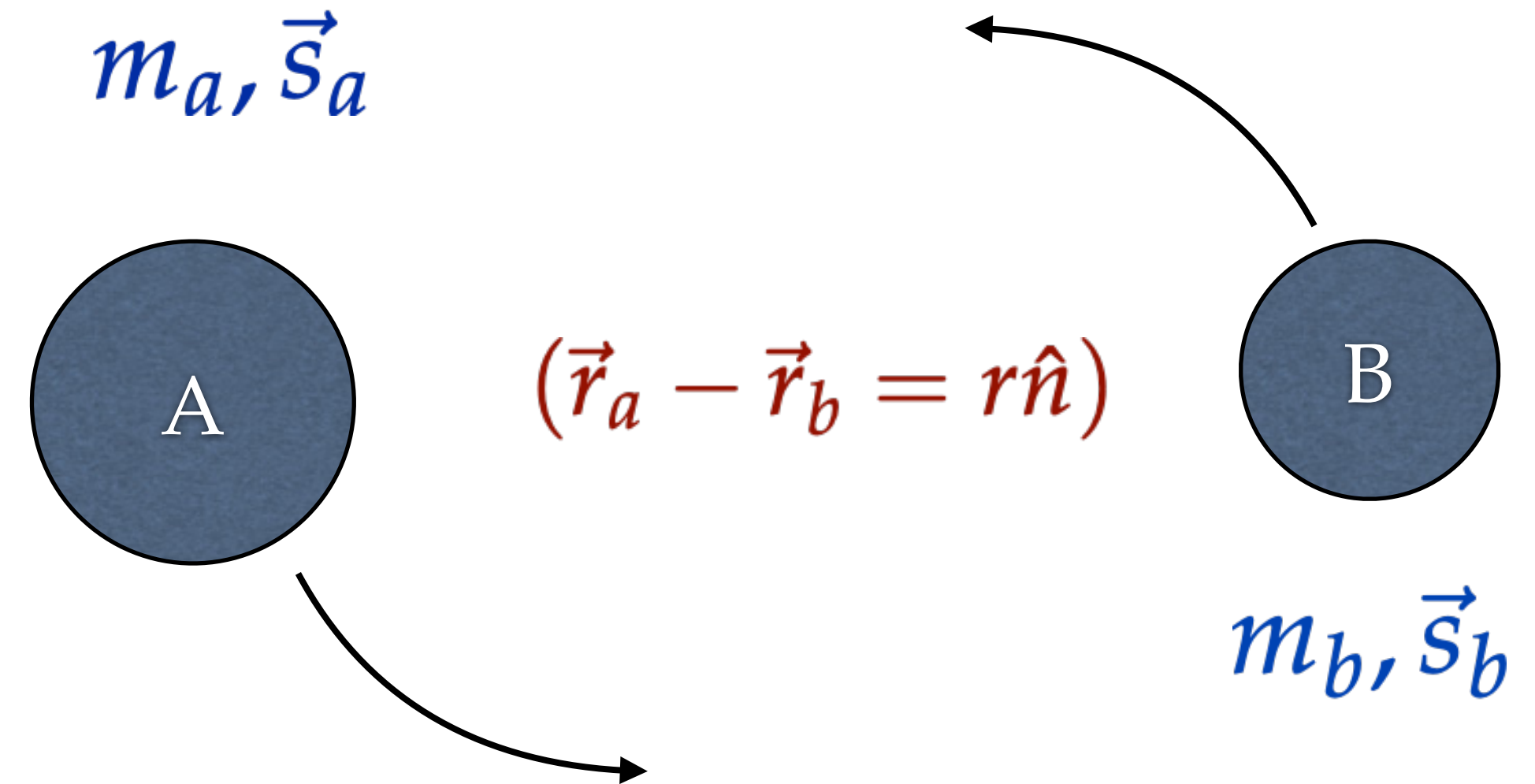
More info to be available in

[H. Lee, SL, S. Mazumdar]

in the 2nd half of 2022

PART II

Binary system with spin



Newton :
$$H = \frac{1}{2m_a} \vec{p}_a^2 + \frac{1}{2m_b} \vec{p}_b^2 - \frac{Gm_a m_b}{r} .$$

Einstein :
$$\mathcal{S} = \frac{1}{16\pi G} \int d^4x \sqrt{-g} R + \mathcal{S}_{\text{matter}} .$$

Newton $\sim$ leading approximation to Einstein

Between Newton and Einstein

Hamiltonian in the COM frame: $\vec{p}_a = \vec{p} = -\vec{p}_b$

Post-Newtonian (PN) :

$$H = \left(\frac{1}{2m_a} + \frac{1}{2m_b} \right) \vec{p}^2 - \frac{Gm_a m_b}{r} + (\text{corrections})$$

Post-Minkowskian (PM) : $E_{a,b} = \sqrt{m_{a,b}^2 + \vec{p}^2}$

$$H = E_a + E_b - \left(\frac{G}{r} \right) \left(\frac{m_a^2 m_b^2}{E_a E_b} \right) \cosh(2\theta) + (\text{corrections})$$

PN vs PM

(1686) (1938) (1974) (2000)

0PN

1PN

2PN

3PN

$$\left(\frac{Gm}{r}\right) \left(\boxed{1} + \boxed{v^2} + \boxed{v^4} + \boxed{v^6} + v^8 + \dots \right) \quad \text{1PM}$$

$$\left(\frac{Gm}{r}\right)^2 \left(\boxed{1} + \boxed{v^2} + \boxed{v^4} + v^6 + \dots \right) \quad \text{2PM}$$

$$\left(\frac{Gm}{r}\right)^3 \left(\boxed{1} + \boxed{v^2} + \boxed{v^4} + \dots \right) \quad \text{3PM}$$

$$\left(\frac{Gm}{r}\right)^4 \left(\boxed{1} + \boxed{v^2} + \dots \right) \quad \text{4PM}$$

⋮

QFT-inspired methods

[Donoghue 96]

[Goldberger, Rothstein 04]

[Hossein, Ross 08]

[Neill, Rothstein 13]

became main-stream around 2018

Small but not negligible

Not small : $|m_a/m_b| = \mathcal{O}(1)$, $|\vec{s}_a/\vec{s}_b| = \mathcal{O}(1)$

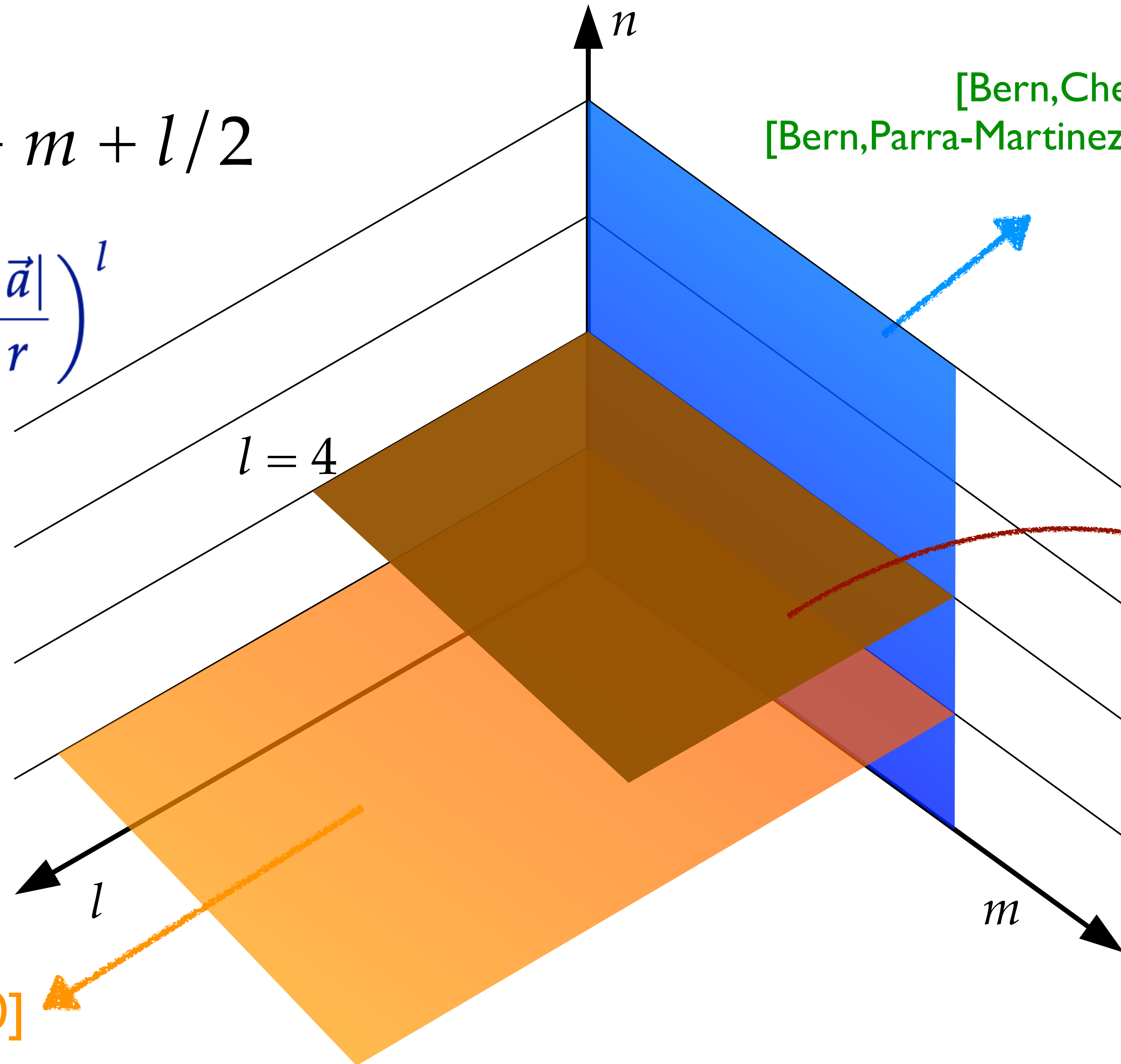
Small : $\frac{Gm}{r}$, $\left(\frac{\vec{a}}{r}\right)^2$ ($\vec{a} \equiv \frac{\vec{s}}{m}$: “spin-length”)
(not negligible)

Possibly small : $\left(\frac{\vec{p}}{m}\right)^2$

Amplitude to GR : the frontier

$$\text{PN order} = n + m + l/2$$

$$\left(\frac{Gm}{r}\right)^n \left(\frac{\vec{p}^2}{m^2}\right)^m \left(\frac{|\vec{a}|}{r}\right)^l$$



[Cheung,Rothstein,Solon 18]

[Bern,Cheung,Roiban,Shen,Solon,Zeng 19]

[Bern,Parra-Martinez,Roiban,Ruf,Shen,Solon,Zeng 21]

[Chen,Chung,Huang,Kim 21]

[Chung,Huang,Kim,SL 20]

Our contribution

Ming-Zhi Chung, Yu-tin Huang, Jung-Wook Kim, SL
[1812.08752][1908.08463][2003.06600]

"Complete" 1PM Hamiltonian

Linear in $\frac{Gm}{r}$

Exact in $\left(\frac{\vec{p}}{m}\right)^2$ and $\left(\frac{\vec{a}}{r}\right)^2$

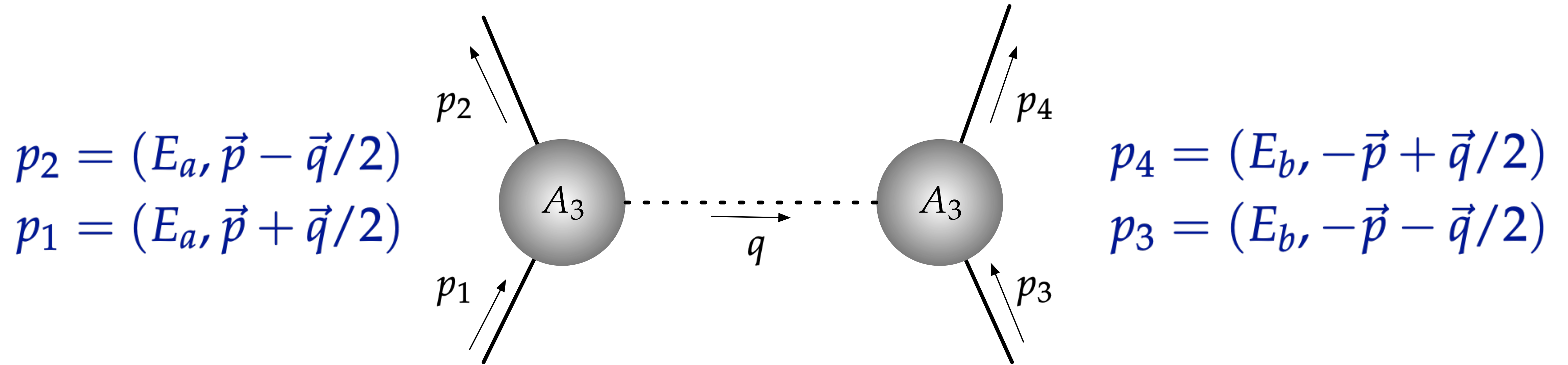
"conservative dynamics"
(ignoring radiation)

Applicable for general spinning body (beyond Kerr BH)

For further progress at 2PM see

[Kim, Levi, Yin 2112.01509]
[Chen, Chung, Huang, Kim 2111.13639]

1PM kinematics



$$E_{a,b} = \sqrt{m_{a,b}^2 + \vec{p}^2}, \quad E = E_a + E_b, \quad \cosh \theta = u_a \cdot u_b = \frac{p_a \cdot p_b}{m_a m_b}$$

Spin-less 1PM potential:

[Cheung, Rothstein, Solon 18]

[Bern, Cheung, Roiban, Shen, Solon, Zeng 19]

$$V_{1\text{PM}}^{(0)} = - \left(\frac{G}{r} \right) \frac{m_a^2 m_b^2}{E_a E_b} \cosh(2\theta)$$

1PM potential with spin

"spin kernel"

$$V_{\text{1PM}} = -\frac{m_a^2 m_b^2}{E_a E_b} \int \frac{d^3 \vec{q}}{(2\pi)^3} \left(\frac{4\pi}{\vec{q}^2} \right) K(\vec{q}, \vec{p}, \vec{a}_{a,b} \cdot \vec{m}_{a,b})$$

$$K = A \times R \quad \text{"(amplitude) x (Thomas precession)"}$$

Ex) Kerr BH

$$A = \frac{1}{2} \left[e^{2\theta} W_a(\tau_a) W_b(\tau_b) + e^{-2\theta} W_a(-\tau_a) W_b(-\tau_b) \right]$$

$$W(x) = e^x$$

$$R = e^{\tilde{\tau}_a} e^{\tilde{\tau}_b}$$

"spin multipole factors"

$$\tau_a = \frac{i}{\sinh \theta} \varepsilon(q, u_a, u_b, a_a), \quad \tilde{\tau}_a = -i \left(\frac{m_b}{r_a E} \right) \varepsilon(q, u_a, u_b, a_a), \quad r_a \equiv 1 + \frac{E_a}{m_a}$$

From Amplitude to Potential

Extract classical contributions from GR amplitudes

[Holstein, Ross 08]
[Neill, Rothstein 13]

Full theory vs PM-EFT

Cheung, Rothstein, Solon [1808.02489]

Spin-multipole moments

Levi, Steinhoff [1501.04956]
Arkani-Hamed, Huang, Huang [1709.04891]

Spin-multipole moments : C and c

Amplitude :

$$\mathcal{A}_3^{+h} = \kappa_h m x^h \left(\frac{\langle \mathbf{12} \rangle^{2S}}{m^{2S}} + c_1 x \frac{\langle \mathbf{12} \rangle^{2S-1} \langle \mathbf{13} \rangle \langle \mathbf{32} \rangle}{m^{2S+1}} + \dots \right) .$$

Potential :

$$\begin{aligned} L_{\text{SI}} = & \sum_{n=1}^{\infty} \frac{(-1)^n}{(2n)!} \frac{C_{2n}}{m^{2n-1}} D_{\mu_{2n}} \cdots D_{\mu_3} \frac{E_{\mu_1 \mu_2}}{\sqrt{u^2}} S^{\mu_1} \cdots S^{\mu_{2n}} \\ & + \sum_{n=1}^{\infty} \frac{(-1)^n}{(2n+1)!} \frac{C_{2n+1}}{m^{2n}} D_{\mu_{2n+1}} \cdots D_{\mu_3} \frac{B_{\mu_1 \mu_2}}{\sqrt{u^2}} S^{\mu_1} \cdots S^{\mu_{2n+1}} . \end{aligned}$$

In the large S limit, the generating functions satisfy

$$A(x) \equiv \sum_{n=0}^{\infty} c_n (x/S)^n , \quad W(x) \equiv \sum_{n=0}^{\infty} \frac{C_n}{n!} x^n \quad \Longrightarrow \quad W(x) = e^x A(x) .$$

BH = minimally coupled particle !

[Hansen 74]
[Steinhardt, Vines 16]
[Vines, 1709.06016]

Kerr BH solution to all orders in spin

$$\bar{h}^{\mu\nu} = \left(u^\mu u^\nu \cos(a \cdot \partial) + u^{(\mu} \epsilon^{\nu)}_{\rho\sigma\gamma} u^\rho a^\sigma \partial^\gamma \frac{\sin(a \cdot \partial)}{(a \cdot \partial)} \right) \frac{4GM}{r}$$

$$\text{Kerr BH : } C_{2n} = C_{2n+1} = 1$$

$$(a^\mu = S^\mu / m)$$

In terms of on-shell 3-point coupling,

$$A(x) = 1, \quad W(x) = e^x$$

$$\text{Kerr BH : } c_0 = 1, \quad c_n = 0 \quad (n \geq 1)$$

1PM parts of PN computations

	LO	NLO	NNLO	N ³ LO	
S^1	✓	✓	✓		★ [Levi, McLeod, von Hippel 2003]
S^2	✓	✓	✓	—	★ [Levi, McLeod, von Hippel 2003]
S^3	✓	★	—	—	
S^4	✓	—	—	—	

★ [Levi, Teng 2008]

[Levi, Mougialakos, Vieira 1912]

A canonical transformation is needed
to compare results in different coordinate choices.

We work in the "isotropic gauge".

No dependence on $(\vec{p} \cdot \hat{n})$ or $(\vec{p} \cdot \vec{a})$
Factorization of the spin-dependence.

Linear in spin

$$V_{S_a^1 S_b^0} = \left(\frac{G}{r^2} \right) m_b [\hat{n} \cdot (\vec{p} \times \vec{a}_a)] X_{(1,0)} .$$

$$X_{(1,0)}^{\text{LO}} = \frac{4m_a + 3m_b}{2m_b} .$$

[Tulczyjew 59]

[Damour 82]

$$X_{(1,0)}^{\text{NLO}} = \left(\frac{18m_a^2 + 8m_a m_b - 5m_b^2}{8m_a^2 m_b^2} \right) \vec{p}^2 .$$

[Tagoshi, Ohashi, Owen 00]

[Faye, Blanchet, Buonanno 06]

[Damour, Jaranowski, Schäfer 07]

$$X_{(1,0)}^{\text{N}^2\text{LO}} = \left(\frac{-15m_a^4 - 15m_a^2 m_b^2 - 12m_a m_b^3 + 7m_b^4}{16m_a^4 m_b^4} \right) \vec{p}^4 .$$

[Hartung, Steinhoff 11] [Levi 11]

[Marsat, Bohe, Faye, Blanchet 12]

[Hartung, Steinhoff, Schäfer 13]

[Levi, Steinhoff 15]

$$X_{(1,0)}^{\text{N}^3\text{LO}} = \left(\frac{84m_a^6 + 50m_a^4 m_b^2 + 84m_a^2 m_b^4 + 80m_a m_b^5 - 45m_b^6}{128m_a^6 m_b^6} \right) \vec{p}^6 .$$

[NEW]

Quadratic in spin

$$V_{S_a^2 S_b^0} = - \left(\frac{G}{r^3} \right) m_a m_b \left[\vec{a}_a^2 - 3(\vec{a}_a \cdot \hat{n})^2 \right] X_{(2,0)} .$$

$$X_{(2,0)}^{\text{LO}} = \frac{1}{2} C_2^{(a)} .$$

[D'Eath 75]
[Thorne, Hartle 85]

$$X_{(2,0)}^{\text{NLO}} = \frac{C_2^{(a)} (6m_a^2 + 16m_a m_b + 6m_b^2) - (8m_a + 7m_b) m_b}{8m_a^2 m_b^2} \vec{p}^2 .$$

[Porto, Rothstein 08][Steinhoff, Hergt, Schäfer 08]

$$X_{(2,0)}^{\text{N}^2\text{LO}} = \frac{C_2^{(a)} (-5m_a^4 + 18m_a^2 m_b^2 - 5m_b^4) - 3(7m_a^2 + 4m_a m_b - 2m_b^2) m_b^2}{16m_a^4 m_b^4} \vec{p}^4 .$$

[Levi 11][Levi, Steinhoff 15-16]

Cubic in spin

$$V_{S_a^3 S_b^0}^{\text{LO}} = \left(\frac{G}{r^4} \right) m_b [\hat{n} \cdot (\vec{p} \times \vec{a}_a)] \left[\vec{a}_a^2 - 5(\vec{a}_a \cdot \hat{n})^2 \right] X_{(3,0)}.$$

$$X_{(3,0)}^{\text{LO}} = \left[C_3^{(a)} \left(\frac{m_a + m_b}{m_b} \right) - \frac{3}{4} C_2^{(a)} \right].$$

[Hergt,Schäfer 07]
[Levi,Steinhoff 14]

$$X_{(3,0)}^{\text{NLO}} = \frac{24 C_3^{(a)} m_a (m_a + m_b) - 3 C_2^{(a)} (6 m_a^2 + 16 m_a m_b + 5 m_b^2) + (12 m_a + 11 m_b) m_b}{16 m_a^2 m_b^2} \vec{p}^2.$$

[Levi,Mougiakakos,Vieira 19]*



$$\begin{aligned} \text{Fig. 1(a7)} = & \frac{1}{2} C_{1(BS3)} \frac{G}{r^4} \frac{m_2}{m_1^2} \left[\vec{S}_1 \cdot \vec{v}_1 \times \vec{v}_2 \left(\vec{S}_1^2 \vec{v}_2 \cdot \vec{n} + 2 \vec{S}_1 \cdot \vec{v}_2 \vec{S}_1 \cdot \vec{n} - 5 (\vec{S}_1 \cdot \vec{n})^2 \vec{v}_2 \cdot \vec{n} \right) \right. \\ & - \vec{S}_1 \cdot \vec{v}_2 \times \vec{n} \left(\vec{S}_1^2 (\vec{v}_1 \cdot \vec{v}_2 - 5 \vec{v}_1 \cdot \vec{n} \vec{v}_2 \cdot \vec{n}) + 2 \vec{S}_1 \cdot \vec{v}_1 \vec{S}_1 \cdot \vec{v}_2 \right. \\ & - 10 \vec{S}_1 \cdot \vec{v}_1 \vec{S}_1 \cdot \vec{n} \vec{v}_2 \cdot \vec{n} - 10 \vec{S}_1 \cdot \vec{v}_2 \vec{S}_1 \cdot \vec{n} \vec{v}_1 \cdot \vec{n} \\ & \left. \left. - 5 (\vec{S}_1 \cdot \vec{n})^2 (\vec{v}_1 \cdot \vec{v}_2 - 7 \vec{v}_1 \cdot \vec{n} \vec{v}_2 \cdot \vec{n}) \right) \right] \end{aligned}$$

$$\begin{aligned} \text{Fig. 1 (a9)} = & -\frac{3}{2}C_1(ES^2)\frac{G}{r^3m_1}\left[2\vec{S}_1\cdot\vec{S}_2\times\vec{v}_2\left(\vec{S}_1\cdot\vec{v}_1\vec{v}_2\cdot\vec{n}+\vec{S}_1\cdot\vec{v}_2\vec{v}_1\cdot\vec{n}\right.\right. \\ & +\vec{S}_1\cdot\vec{n}\vec{v}_1\cdot\vec{v}_2-5\vec{S}_1\cdot\vec{n}\vec{v}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n}\left.\right)+\vec{S}_2\cdot\vec{v}_1\times\vec{v}_2\left(S_2^2\vec{v}_2\cdot\vec{n}-2\vec{S}_1\cdot\vec{v}_2\vec{S}_1\cdot\vec{n}\right. \\ & +5(\vec{S}_1\cdot\vec{n})^2\vec{v}_2\cdot\vec{n}-\vec{S}_2\cdot\vec{v}_2\times\vec{n}\left(S_1^2(\vec{v}_1\cdot\vec{v}_2-5\vec{v}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n})-2\vec{S}_1\cdot\vec{v}_1\vec{S}_1\cdot\vec{v}_2\right. \\ & +10\vec{S}_1\cdot\vec{v}_1\vec{S}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n}+10\vec{S}_1\cdot\vec{v}_2\vec{S}_1\cdot\vec{n}\vec{v}_1\cdot\vec{n}\left.\right) \\ & \left.+5(\vec{S}_1\cdot\vec{n})^2(\vec{v}_1\cdot\vec{v}_2-7\vec{v}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n})\right] \\ & +\frac{1}{2}C_1(ES^2)\frac{G}{r^3m_1}\left[2\vec{S}_1\cdot\vec{S}_2\times\vec{v}_2(\vec{S}_1\cdot\vec{v}_2-3\vec{S}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n})\right. \\ & -2(\vec{S}_1\cdot\vec{S}_2\times\vec{a}_2+\vec{S}_1\cdot\vec{\dot{S}}_2\times\vec{v}_2)(\vec{S}_1\cdot\vec{v}_1-3\vec{S}_1\cdot\vec{n}\vec{v}_1\cdot\vec{n}) \\ & +2\vec{S}_1\cdot\vec{S}_2\times\vec{v}_2(\vec{S}_1\cdot\vec{v}_2-3\vec{S}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n}) \\ & -(\vec{S}_2\cdot\vec{v}_1\times\vec{a}_2+\vec{\dot{S}}_2\cdot\vec{v}_1\times\vec{v}_2)(S_1^2+3(\vec{S}_1\cdot\vec{n})^2)+6\vec{S}_2\cdot\vec{v}_2\times\vec{n}(\vec{S}_1\cdot\vec{S}_1\vec{v}_2\cdot\vec{n} \\ & -\vec{S}_1\cdot\vec{v}_2\vec{S}_1\cdot\vec{n}-\vec{S}_1\cdot\vec{v}_2\vec{\dot{S}}_1\cdot\vec{n}+5\vec{S}_1\cdot\vec{n}\vec{S}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n}) \\ & \left.-3(\vec{\dot{S}}_2\cdot\vec{v}_2\times\vec{n}+\vec{S}_2\cdot\vec{a}_2\times\vec{n})(S_1^2\vec{v}_1\cdot\vec{n}-2\vec{S}_1\cdot\vec{v}_1\vec{S}_1\cdot\vec{n}+5(\vec{S}_1\cdot\vec{n})^2\vec{v}_1\cdot\vec{n})\right] \end{aligned}$$

$$\begin{aligned} \text{Fig. 1(a10)} = & \frac{3}{2}C_{1(ES^2)}\frac{G}{r^4}\frac{1}{m_1}\left[2\vec{S}_1\cdot\vec{S}_2\times\vec{v}_1\left(\dot{\vec{S}}_1\cdot\vec{v}_1\vec{v}_2\cdot\vec{n}+\vec{S}_1\cdot\vec{v}_2\vec{v}_1\cdot\vec{n}+\vec{S}_1\cdot\vec{n}\vec{v}_1\cdot\vec{v}_2\right.\right. \\ & -5\vec{S}_1\cdot\vec{n}\vec{v}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n})-\vec{S}_2\cdot\vec{v}_1\times\vec{v}_2\left(S_1^2\vec{v}_1\cdot\vec{n}-2\vec{S}_1\cdot\vec{v}_1\vec{S}_1\cdot\vec{n}\right. \\ & +5(\vec{S}_1\cdot\vec{n})^2\vec{v}_1\cdot\vec{n})-\vec{S}_2\cdot\vec{v}_1\times\vec{n}\left(\vec{S}_1^2(\vec{v}_1\cdot\vec{v}_2-5\vec{v}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n})\right. \\ & -2\vec{S}_1\cdot\vec{v}_1\vec{S}_1\cdot\vec{v}_2+10\vec{S}_1\cdot\vec{v}_1\vec{S}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n}+10\vec{S}_1\cdot\vec{v}_2\vec{S}_1\cdot\vec{n}\vec{v}_1\cdot\vec{n} \\ & \left.\left.+5(\vec{S}_1\cdot\vec{n})^2\left(\vec{v}_1\cdot\vec{v}_2+7\vec{v}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n}\right)\right)\right] \\ & -\frac{1}{2}C_{1(ES^2)}\frac{G}{r^3}\frac{1}{m_1}\left[\vec{S}_1\cdot\vec{S}_2\times\vec{v}_1(\dot{\vec{S}}_1\cdot\vec{v}_2-3\dot{\vec{S}}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n})\right. \\ & +2\vec{S}_1\cdot\vec{S}_2\times\vec{a}_1(\vec{S}_1\cdot\vec{v}_2-3\dot{\vec{S}}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n})+\dot{\vec{S}}_1\cdot\vec{S}_2\times\vec{v}_1(\vec{S}_1\cdot\vec{v}_2-3\vec{S}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n}) \\ & -2\vec{S}_1\cdot\dot{\vec{S}}_2\times\vec{v}_1(\vec{S}_1\cdot\vec{v}_1-3\dot{\vec{S}}_1\cdot\vec{n}\vec{v}_1\cdot\vec{n})-\vec{S}_1\cdot\vec{S}_2\times\vec{v}_2(\dot{\vec{S}}_1\cdot\vec{v}_1-3\dot{\vec{S}}_1\cdot\vec{n}\vec{v}_1\cdot\vec{n}) \\ & -\dot{\vec{S}}_1\cdot\vec{S}_2\times\vec{v}_2(\vec{S}_1\cdot\vec{v}_1-3\dot{\vec{S}}_1\cdot\vec{n}\vec{v}_1\cdot\vec{n})+3\vec{S}_1\cdot\vec{S}_2\times\vec{n}\left(\dot{\vec{S}}_1\cdot\vec{v}_1\vec{v}_2\cdot\vec{n}\right. \\ & \left.+\dot{\vec{S}}_1\cdot\vec{v}_2\vec{v}_1\cdot\vec{n}+\dot{\vec{S}}_1\cdot\vec{n}\vec{v}_1\cdot\vec{v}_2-5\dot{\vec{S}}_1\cdot\vec{n}\vec{v}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n})\right) \\ & +3\dot{\vec{S}}_1\cdot\vec{S}_2\times\vec{n}(\vec{S}_1\cdot\vec{v}_1\vec{v}_2\cdot\vec{n}+\vec{S}_1\cdot\vec{v}_2\vec{v}_1\cdot\vec{n}+\vec{S}_1\cdot\vec{n}\vec{v}_1\cdot\vec{v}_2 \\ & -5\vec{S}_1\cdot\vec{n}\vec{v}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n})-2\dot{\vec{S}}_2\cdot\vec{v}_1\times\vec{v}_2(\dot{\vec{S}}_1\cdot\vec{S}_1+3\dot{\vec{S}}_1\cdot\vec{n}\vec{S}_1\cdot\vec{n}) \\ & -\vec{S}_2\cdot\vec{a}_1\times\vec{v}_2\left(S_1^2+3(\vec{S}_1\cdot\vec{n})^2\right)+6\dot{\vec{S}}_2\cdot\vec{v}_1\times\vec{n}\left(\dot{\vec{S}}_1\cdot\vec{S}_1\vec{v}_2\cdot\vec{n}-\vec{S}_1\cdot\vec{v}_2\dot{\vec{S}}_1\cdot\vec{n}\right. \\ & \left.-\dot{\vec{S}}_1\cdot\vec{v}_2\vec{S}_1\cdot\vec{n}+5\dot{\vec{S}}_1\cdot\vec{n}\vec{S}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n})+3\dot{\vec{S}}_2\cdot\vec{a}_1\times\vec{n}\left(S_1^2\vec{v}_2\cdot\vec{n}-2\vec{S}_1\cdot\vec{v}_2\vec{S}_1\cdot\vec{n}\right. \\ & \left.+5(\vec{S}_1\cdot\vec{n})^2\vec{v}_2\cdot\vec{n})+3\dot{\vec{S}}_2\cdot\vec{v}_1\times\vec{n}(3\dot{\vec{S}}_1^2\vec{v}_1\cdot\vec{n}+2\vec{S}_1\cdot\vec{v}_1\vec{S}_1\cdot\vec{n}-5(\vec{S}_1\cdot\vec{n})^2\vec{v}_1\cdot\vec{n})\right] \\ & +\frac{1}{2}C_{1(ES^2)}\frac{G}{r^2}\frac{1}{m_1}\left[\vec{S}_1\cdot\dot{\vec{S}}_2\times\vec{v}_1\dot{\vec{S}}_1\cdot\vec{n}+\dot{\vec{S}}_1\cdot\dot{\vec{S}}_2\times\vec{v}_1\vec{S}_1\cdot\vec{n}\right. \\ & +\vec{S}_1\cdot\vec{S}_2\times\vec{v}_2\dot{\vec{S}}_1\cdot\vec{n}+2\dot{\vec{S}}_1\cdot\vec{S}_2\times\vec{v}_2\dot{\vec{S}}_1\cdot\vec{n}+\dot{\vec{S}}_1\cdot\vec{S}_2\times\vec{v}_2\vec{S}_1\cdot\vec{n} \\ & +\vec{S}_1\cdot\vec{S}_2\times\vec{n}(\dot{\vec{S}}_1\cdot\vec{v}_2-3\dot{\vec{S}}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n})-\vec{S}_1\cdot\dot{\vec{S}}_2\times\vec{n}(\dot{\vec{S}}_1\cdot\vec{v}_1-3\dot{\vec{S}}_1\cdot\vec{n}\vec{v}_1\cdot\vec{n}) \\ & +2\dot{\vec{S}}_1\cdot\vec{S}_2\times\vec{n}(\dot{\vec{S}}_1\cdot\vec{v}_2-3\dot{\vec{S}}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n})-\dot{\vec{S}}_1\cdot\dot{\vec{S}}_2\times\vec{n}(\vec{S}_1\cdot\vec{v}_1-3\vec{S}_1\cdot\vec{n}\vec{v}_1\cdot\vec{n}) \\ & +\dot{\vec{S}}_1\cdot\vec{S}_2\times\vec{n}(\vec{S}_1\cdot\vec{v}_2-3\vec{S}_1\cdot\vec{n}\vec{v}_2\cdot\vec{n})+2\vec{S}_1\cdot\dot{\vec{S}}_2\times\vec{a}_1\vec{S}_1\cdot\vec{n} \\ & \left.+2\dot{\vec{S}}_2\cdot\vec{v}_1\times\vec{n}(\dot{\vec{S}}_1\cdot\vec{S}_1+3\dot{\vec{S}}_1\cdot\vec{n}\vec{S}_1\cdot\vec{n})+\dot{\vec{S}}_2\cdot\vec{a}_1\times\vec{n}\left(S_1^2+3(\vec{S}_1\cdot\vec{n})^2\right)\right] \\ & -\frac{1}{2}C_{1(ES^2)}\frac{G}{r}\frac{1}{m_1}\left[\vec{S}_1\cdot\dot{\vec{S}}_2\times\vec{n}\dot{\vec{S}}_1\cdot\vec{n}+2\dot{\vec{S}}_1\cdot\dot{\vec{S}}_2\times\vec{n}\dot{\vec{S}}_1\cdot\vec{n}+\dot{\vec{S}}_1\cdot\dot{\vec{S}}_2\times\vec{n}\vec{S}_1\cdot\vec{n}\right]. \end{aligned}$$

PART III

[Hanson, Regge 1974]

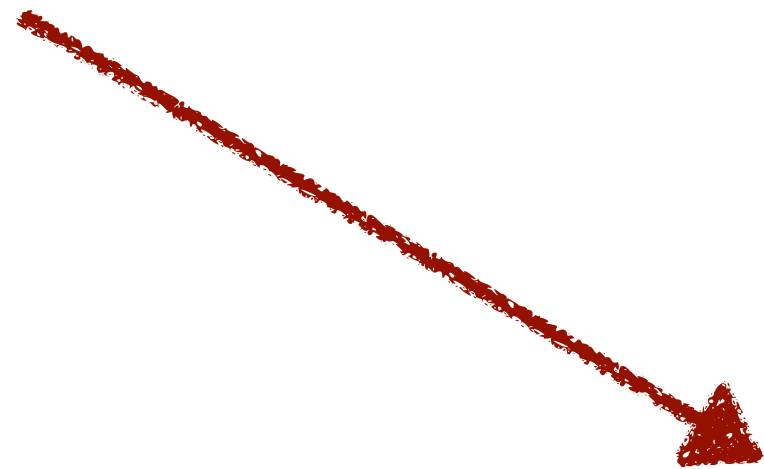
[Fedoruk, Lukierski 2014]

The Relativistic Spherical Top as a Massive Twistor

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"Spherical Top" in special relativity

$$S = \int dt \left(\frac{1}{2} m \vec{v}^2 \right) \xrightarrow{\text{"spin"}} S = \int dt \left(\frac{1}{2} m \vec{v}^2 + \frac{1}{2} I \vec{\omega}^2 \right)$$

(1+3)-dim
space-time



$$S = - \int ds \sqrt{-\eta_{ab} \dot{x}^a \dot{x}^b} \longrightarrow$$

$$S = (?)$$

[Frenkel 1926][Thomas 1926]
... [Hanson, Regge 1974]

3 translation d.o.f. + 3 rotation d.o.f.

$$x, p \in \mathbb{R}^{1,3}$$

$$\Lambda \in \text{SO}(1,3)$$

$$S \in \mathfrak{so}(1,3)$$

$$\{x^\mu, p_\nu\} = \delta^\mu_\nu,$$

$$\{\Lambda^\rho_A, S_{\mu\nu}\} = -(\delta^\rho_\mu \Lambda_{\nu A} - \delta^\rho_\nu \Lambda_{\mu A}),$$

$$\{S_{\mu\nu}, S_{\rho\sigma}\} = -(\eta_{\nu\rho} S_{\mu\sigma} - \eta_{\mu\rho} S_{\nu\sigma} - \eta_{\nu\sigma} S_{\mu\rho} + \eta_{\mu\sigma} S_{\nu\rho}).$$

Translation: $4 \rightarrow 3$

$$p^2 + m^2 = 0$$

"mass shell constraint"

Rotation: $6 \rightarrow 3$

(?)

"spin constraints"

Spin constraints

Intuition : remove the "boosts" from Λ^μ_A .

But, boosts do not form a closed subalgebra.

$$\Lambda^\mu_0 S_{\mu\nu} = 0 \quad (\text{temporal})$$

$$\hat{p}^\mu S_{\mu\nu} = 0 \quad (\text{covariant})$$

$$(\hat{p}^\mu + \Lambda^\mu_0) S_{\mu\nu} = 0 \quad (\text{Pryce-Newton-Wigner})$$

[Pryce 1948][Newton,Wigner 1949]

Our proposal

(mass shell)

$$\phi_0 = \frac{1}{2}(p^2 + m^2),$$

$$\chi^0 = \frac{1}{p^2} x^\mu p_\mu,$$

(modified PNW)

$$\phi_a = \frac{1}{2}(\hat{p}^\mu + \Lambda^\mu{}_0) S_{\mu\nu} \Lambda^\nu{}_a,$$

$$\chi^a = \hat{p}_\mu \Lambda^{\mu a}.$$

[Steinhoff 2015]

Canonically conjugate pairs of

"gauge generators" and "gauge-fixing conditions"

$$\{\chi^A, \phi_B\} \approx \delta^A{}_B$$

$$\{\phi_A, \phi_B\} \approx 0$$

Dirac bracket

$$\{f, g\}_* = \{f, g\} - \{f, \phi_A\}\{\chi^A, g\} + \{f, \chi^A\}\{\phi_A, g\}$$

The only dynamical information
not governed by symmetry is ...

the "Regge trajectory"

$$p^2 + m^2(\tilde{S}^2) = 0$$

[Hanson, Regge 1974]

Equations of motion

Lagrange multipliers

$$\begin{aligned} \mathcal{S} &= \int d\sigma \left(p_\mu \dot{x}^\mu + \frac{1}{2} S_{\mu\nu} \Omega^{\mu\nu} - \kappa \phi_0 - \kappa^a \phi_a \right) \\ &= \int d\sigma \left(p_\mu \dot{x}^\mu + \frac{1}{2} S_{\mu\nu} \Omega^{\mu\nu} - \frac{\kappa}{2} (p^2 + m^2) - \frac{\kappa^a}{2} (\hat{p}^\mu + \Lambda^\mu{}_0) S_{\mu\nu} \Lambda^\nu{}_a \right) . \end{aligned}$$

$$\begin{aligned} \frac{dp_\mu}{d\tau} &= 0, \quad p^\mu = m \frac{dx^\mu}{d\tau}, \\ \frac{dS_{\mu\nu}}{d\tau} &= 0, \quad S^{\mu\nu} = \frac{1}{2m'} \bar{\Omega}^{\mu\nu}. \end{aligned}$$

$$\left(\bar{\Omega}^{\mu\nu} = \Lambda^{\mu A} \frac{d\Lambda^\nu{}_A}{d\tau} \right)$$

Massive twistor

Twistor coordinates :

$$(8_{\mathbb{C}} = 16_{\mathbb{R}})$$

$$Z_A^I = \begin{pmatrix} \lambda_\alpha^I \\ i\mu^{\dot{\alpha}I} \end{pmatrix}, \quad \bar{Z}_I^B = (Z_A^I)^\dagger A^{\bar{A}B} = \begin{pmatrix} -i\bar{\mu}_I^\beta & \bar{\lambda}_{I\dot{\beta}} \end{pmatrix}.$$

conformal symmetry
to be broken by
the mass shell condition

$$\text{SU}(2,2) \times \text{SU}(2)$$

interpreted as the
"little group" symmetry
of a massive particle

Poisson bracket algebra :

$$\{Z_A^I, \bar{Z}_J^B\} = -i \delta_A^B \delta^I_J \implies \{\lambda_\alpha^I, \bar{\mu}_J^\beta\} = \delta_\alpha^\beta \delta^I_J, \quad \{\bar{\lambda}_{I\dot{\alpha}}, \mu^{\dot{\beta}J}\} = \delta_{\dot{\alpha}}^{\dot{\beta}} \delta_I^J.$$

Constraints

$$\phi_0 = \frac{1}{2}(p^2 + m^2), \quad \chi^0 = \frac{1}{p^2} x^\mu p_\mu.$$

Reduction from 16 to 12.

Spin constraints are **not** needed.

Mass-shell constraint should be "complexified".

$$\begin{aligned} \phi &= -\frac{1}{2} \left(\det(\lambda) - m(\tilde{S}^2) \right), & \bar{\chi} &= \frac{1}{\det(\lambda)} \langle \bar{\mu} \lambda \rangle, \\ \bar{\phi} &= -\frac{1}{2} \left(\det(\bar{\lambda}) - m(\tilde{S}^2) \right), & \chi &= \frac{1}{\det(\bar{\lambda})} [\bar{\lambda} \mu]. \end{aligned}$$

Twistor model "solves" the spin constraints.

Map to the spherical top

$$p_{\alpha\dot{\alpha}} = (\sigma^\mu)_{\alpha\dot{\alpha}} p_\mu = \begin{pmatrix} p_0 + p_3 & p_1 - ip_2 \\ p_1 - ip_2 & p_0 - p_3 \end{pmatrix}$$

The map (assuming constraints)

$$p_{\alpha\dot{\alpha}} = -\lambda_\alpha{}^I \bar{\lambda}_{I\dot{\alpha}} , \quad x^{\dot{\alpha}\alpha} = \frac{1}{2} \left(\frac{1}{\det(\lambda)} \mu^{\dot{\alpha}I} \lambda^{\alpha J} \epsilon_{IJ} + \frac{1}{\det(\bar{\lambda})} \epsilon^{IJ} \bar{\lambda}_I{}^{\dot{\alpha}} \bar{\mu}_J{}^\alpha \right) ,$$

$$(\Lambda_{\alpha\dot{\alpha}})^I{}_J = \frac{2\lambda_\alpha{}^I \bar{\lambda}_{J\dot{\alpha}}}{|\det(\lambda)|} , \quad S_{\alpha\beta} = \frac{1}{2} \left(\bar{\mu}_{I(\alpha} \lambda_{\beta)}^I - \frac{1}{\det(\lambda)} \lambda_{(\alpha}{}^I \lambda_{\beta)}^J (\mu^{\dot{\gamma}}{}_I \bar{\lambda}_{J\dot{\gamma}}) \right) .$$

incidence relation
complex Minkowski space

$$\mu^{\dot{\alpha}I} = -z^{\dot{\alpha}\beta} \lambda_\beta{}^I , \quad z^\mu = x^\mu + iy^\mu .$$

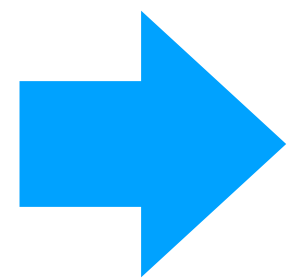
Hints

$$J_{\mu\nu} = x_\mu p_\nu - x_\nu p_\mu + S_{\mu\nu} , \quad S_{\mu\nu} = \frac{1}{2} \epsilon_{\mu\nu}{}^{\rho\sigma} (y_\rho p_\sigma - y_\sigma p_\rho) .$$

Equations of motion

$$\phi = -\frac{1}{2} \left(\det(\lambda) - m(\tilde{S}^2) \right) ,$$
$$\bar{\phi} = -\frac{1}{2} \left(\det(\bar{\lambda}) - m(\tilde{S}^2) \right) .$$

$$\mathcal{S} = \int d\sigma \left(\frac{i}{2} (\bar{Z}_I^A \dot{Z}_A^I - Z_A^I \dot{\bar{Z}}_I^A) - \bar{\kappa} \phi - \kappa \bar{\phi} \right)$$



$$\dot{\lambda}_\alpha^I = -i(\operatorname{Re} \kappa) m' \lambda_\alpha^J W_J^I ,$$

$$\dot{\mu}^{\dot{\alpha}I} = -i(\operatorname{Re} \kappa) m' \mu^{\dot{\alpha}J} W_J^I + \frac{\kappa}{2} \bar{\lambda}^{\dot{\alpha}I} .$$

Perfect agreement with the spherical top model !

Quantization

$$\bar{\mu}_I{}^\alpha = -i \frac{\partial}{\partial \lambda_\alpha^I} , \quad \mu^{\dot{\alpha}I} = -i \frac{\partial}{\partial \bar{\lambda}_{I\dot{\alpha}}} .$$

Quantize first, and then impose constraints

On-shell fields and Fierz-Pauli / Bargmann-Wigner equation:

$$\begin{aligned} \Phi_{\alpha_1 \alpha_2 \dots \alpha_{2s}}^{(2s,0)}(x) &= c_{I_1 \dots I_{2s}} \int d[\lambda, \bar{\lambda}] \lambda_{\alpha_1}^{I_1} \lambda_{\alpha_2}^{I_2} \dots \lambda_{\alpha_{2s}}^{I_{2s}} \exp \left(i x^{\dot{\beta}\beta} \bar{\lambda}_{\dot{\beta}J} \lambda_{\beta}^J \right) , \\ \Phi_{\dot{\alpha}_1 | \alpha_2 \dots \alpha_{2s}}^{(2s-1,1)}(x) &= c_{I_1 \dots I_{2s}} \int d[\lambda, \bar{\lambda}] \bar{\lambda}_{\dot{\alpha}_1}^{I_1} \lambda_{\alpha_2}^{I_2} \dots \lambda_{\alpha_{2s}}^{I_{2s}} \exp \left(i x^{\dot{\beta}\beta} \bar{\lambda}_{\dot{\beta}J} \lambda_{\beta}^J \right) , \\ &\vdots \\ \Phi_{\dot{\alpha}_1 \dot{\alpha}_2 \dots \dot{\alpha}_{2s}}^{(0,2s)}(x) &= c_{I_1 \dots I_{2s}} \int d[\lambda, \bar{\lambda}] \bar{\lambda}_{\dot{\alpha}_1}^{I_1} \bar{\lambda}_{\dot{\alpha}_2}^{I_2} \dots \bar{\lambda}_{\dot{\alpha}_{2s}}^{I_{2s}} \exp \left(i x^{\dot{\beta}\beta} \bar{\lambda}_{\dot{\beta}J} \lambda_{\beta}^J \right) , \end{aligned}$$

where $c_{I_1 \dots I_{2s}}$ is a totally symmetrized SU(2) tensor, and the integration measure is

$$d[\lambda, \bar{\lambda}] = d^4 \lambda d^4 \bar{\lambda} \left[\delta(\det(\lambda) - m) \delta(\det(\bar{\lambda}) - m) \right] .$$

Newman-Janis shift

[Newman, Janis 1965]

[Guevara, Maybee, O'Connell, Ochirov, Vines 2012.11570]

Complex Minkowski spacetime with the "spin-length" vector :

$$x^\mu \rightarrow x^\mu \pm iy^\mu, \quad y_\mu = \frac{1}{m^2} W_\mu = \frac{1}{m^2} \varepsilon_{\mu\nu\rho\sigma} p^\nu S^{\rho\sigma}$$

opens a way to couple the twistor model

to background fields.

[J. Kim, SL, work in progress]

Summary and Outlook

Spinor-helicity variables are useful.

- QFT (QED, QCD, ...)
- GR (post-Minkowskian expansion, Newman-Janis shift, etc.)

Perhaps, we should rewrite some chapters of textbooks ?