

Black hole microstates in gauge theories

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Introduction

To better understand quantum aspect of gravity, we want to put gravity in a box

A natural choice is **anti-de Sitter** (AdS) spacetime

- solution of Einstein equation with **negative cosmological constant**

$$ds^2 = - \left(1 + \frac{r^2}{\ell^2}\right) dt^2 + \left(1 + \frac{r^2}{\ell^2}\right)^{-1} dr^2 + r^2 d\Omega_{d-2}^2$$

- gravitating potential well

$$g_{tt} \approx -(1 + 2\Phi) \quad \Rightarrow \quad \Phi \approx \frac{r^2}{2\ell^2}$$

confining particles toward AdS center

- conformal boundary at asymptotic infinity

Black holes in AdS

Schwarzschild black hole (d=5) $ds^2 = - \left(1 + \frac{r^2}{\ell^2} - \frac{\mu}{r^2}\right) dt^2 + \left(1 + \frac{r^2}{\ell^2} - \frac{\mu}{r^2}\right)^{-1} dr^2 + r^2 d\Omega_3^2$

- temperature $T = \frac{r_h}{\pi\ell^2} + \frac{1}{2\pi r_h}, r_h^2 = \frac{-\ell^2 + \sqrt{\ell^2(\ell^2 + 4\mu)}}{2}$

- two branches: small & large black hole

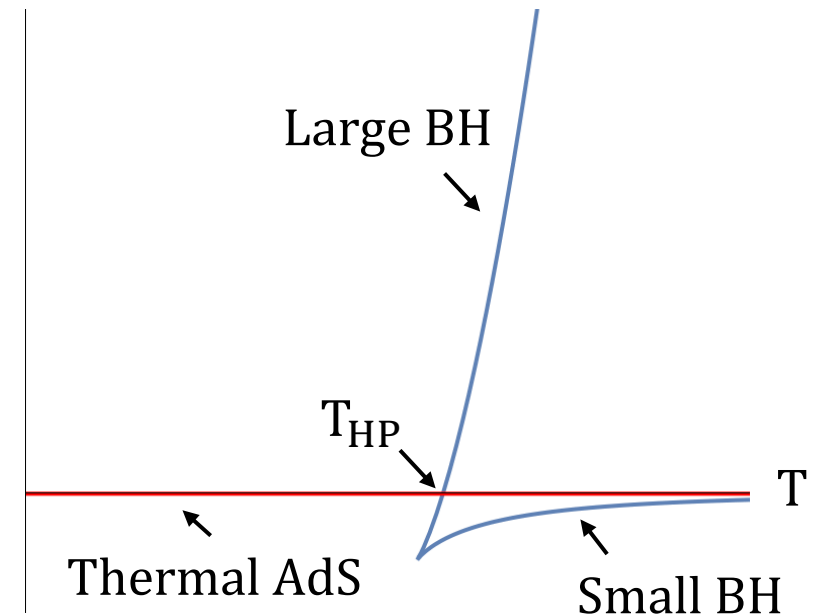
Thermal partition function $Z(\beta) = \int \mathcal{D}g \exp \left[-\frac{1}{G_N} I_{\text{EH}} \right]$

- low T: Gas of gravitons $\log Z \sim O(1)$

Hawking-Page transition ($T_{\text{HP}} = \frac{3}{2\pi\ell}$)

- high T: Large AdS black holes $\log Z \sim O(1/G_N)$

$G_N \cdot \text{Log } Z$



AdS/CFT correspondence

Maldacena proposed [Maldacena '97]

String theory on $\text{AdS}_5 \times S^5 \leftrightarrow \mathcal{N} = 4$ Super-Yang-Mills theory

boundary of AdS
↓

$S^3 \times R$

on

Especially, we can study gravity in AdS by large N , strongly coupled CFT

$$\ell_{\text{AdS}} \gg l_s \quad g_s \rightarrow 0 \quad \Leftrightarrow \quad N \rightarrow \infty \quad \lambda \gg 1 \quad \left(G_N \sim \frac{1}{N^2} \right)$$

Hawking-Page transition can be understood as deconfinement transition in CFT

- low T : Gas of gravitons

$$\log Z \sim O(1)$$

Confinement

↑
Hawking-Page transition

↑
Confinement-deconfinement transition

- high T : Large AdS black holes

$$\log Z \sim O(1/G_N) \\ \sim O(N^2)$$

↓
Deconfined Quark-gluon plasma

Supersymmetric black holes in AdS_5

To be quantitative, we consider **supersymmetric** setup

Supersymmetric black holes in $\text{AdS}_5 \times S^5$ are known [Gutowski-Reall] [Chong-Cvetič-Lu-Pope]
[Kunduri-Luccietti-Reall]

- three electric charges R_1, R_2, R_3 & two angular momenta J_1, J_2
- mass (energy) $E = (R_1 + R_2 + R_3 + J_1 + J_2)$
- preserving **two supersymmetry**

Dual CFT = $\mathcal{N} = 4$ super-Yang-Mills theory with $SU(N)$ gauge group

We consider the operators preserving **two specific supercharges** Q, S

$$[Q, O] = 0 \ \& \ [S, O] = 0 \quad \Leftrightarrow \quad [\{Q, S\}, O] = 0$$

- saturate the BPS bound $E - (R_1 + R_2 + R_3 + J_1 + J_2) = 0$

which follows from the algebra $2\{Q, S\} = E - (R_1 + R_2 + R_3 + J_1 + J_2) \geq 0$

Cohomology problem

It is hard to study the operators in the strong coupling ...

1. we will begin with the **weak coupling** regime

→ assuming a sort of non-renormalization of spectrum

- entropy of BPS BH are counted from (coupling independent) index

2. we will study supercharge cohomology instead of BPS operators

- due to nilpotency of the supercharge $Q^2 = 0$

$$\begin{array}{ccc} [\{Q, Q^\dagger\}, O] = 0 & \overset{\text{1-to-1}}{\Leftrightarrow} & [Q, O] = 0 / (O \sim O + Q - \text{exact}) \\ \text{Harmonic forms} & & Q\text{-cohomology classes} \end{array}$$

Cohomology problem

In the free limit $g_{\text{YM}} \rightarrow 0$, BPS operators can be constructed by free BPS letters

$$\bar{\phi}^m, \psi_{m+}, \bar{\lambda}_{\dot{\alpha}}, f_{++} \equiv F_{1+i2, 3+i4}, \quad \text{and} \quad \partial_{++} \equiv \partial_1 - i\partial_2, \partial_{+\dot{-}} \equiv \partial_3 - i\partial_4$$

i.e. all gauge invariant operators are BPS $\rightarrow$ Too many !

1-loop order correction to anomalous dimension $\sim O(g_{\text{YM}}^2)$

At $g_{\text{YM}} \ll 1$, **half-loop Q** action on elementary fields are

$$Q\bar{\phi}^m = 0, Q\psi_{m+} = -ig_{\text{YM}}\epsilon_{mnp}[\bar{\phi}^n, \bar{\phi}^p], Qf_{++} = -ig_{\text{YM}}[\psi_{m+}, \bar{\phi}^m]$$

$$Q\bar{\lambda}_{\dot{\alpha}} = 0, [Q, D_{+\dot{\alpha}}](\cdots) = -ig_{\text{YM}}[\bar{\lambda}_{\dot{\alpha}}, (\cdots)]$$

$\rightarrow$ many free cohomology get lifted at 1-loop

Among free BPS operators which remain BPS at weak coupling?

Graviton cohomology

There is a well-known class of cohomology called **graviton** [Gunaydin, Marcus '84]
[Kim, Romans, van Nieuwenhuizen '85]
which corresponds to BPS Kaluza-Klein gravitons in $\text{AdS}_5 \times S^5$

- stress-tensor multiplet + KK tower

Explicitly, graviton multiplet consists of

- chiral primary $\text{tr}(\bar{\phi}^{(i_1} \dots \bar{\phi}^{i_n)})$ & superconformal descendants
acting supercharges $Q_+^m, \bar{Q}_{m\dot{\alpha}}$ and BPS derivatives $\partial_{+\dot{\alpha}}$ on chiral primary

Graviton cohomology

- the above single trace gravitons & their multiparticle states
- explains **low E spectrum**
- Q-closed for any N (often called **monotone**)

Trace relations

As we grow the energy $E \sim N$, **finite N effect** called trace relation become important

- relations between single/multi-trace operators (mesons/baryons see the gluons)

e.g. SU(2) SYM $\text{tr}(\bar{\phi}_1^4) = \frac{1}{2} \text{tr}(\bar{\phi}_1^2) \text{tr}(\bar{\phi}_1^2)$, $\text{tr}(\bar{\phi}_1^2) \text{tr}(\bar{\phi}_2^2) - (\text{tr}(\bar{\phi}_1 \bar{\phi}_2))^2 \sim Q \text{tr}(\psi_3 + [\bar{\phi}_1, \bar{\phi}_2])$

→ **reduces the states** c.f. Cayley-Hamilton

In gravity dual, this can be understood as stringy exclusion principle [Maldacena, Strominger '98]

- effect of giant gravitons: D-branes wrapping S^3 which have maximal size
[McGreevy, Susskind, Toumbas '00] [Grisaru, Myers, Tafjord '00]

In interacting supersymmetric theories, trace relations can **generate new operators**

$$Q\mathcal{O} = (\text{Trace relation}) = 0$$

- explains the rest of the cohomology, which we call non-graviton cohomology

- this phenomena is also known as **fortuity** [Chang, Lin '24]

Non-graviton cohomology

The very first example is obtained only recently in SU(2) SYM [Choi-Kim-Lee-Park '22]

- fortuity in many other models are studied then after

[Chang-Chen-Sia-Yang '24] [Chang-Lin-Zhang '25] [Belin-Singh-Vadala-Zaffaroni '25]
[Behan-de Gioia '25] [Kim-JL-Lee-Oh '25] [Seunggyu's talk]

In $SU(3)$ SYM, we constructed the lowest energy cohomology

$$\begin{aligned} -6\mathcal{O}^3 = & 288v^{(j}{}_av^{k)a}{}_i \times \epsilon_{c_1c_2(j} \text{tr} (\bar{\phi}^{c_1} \bar{\phi}^{c_2} \bar{\phi}^i \psi_{k+}) - 72v^a{}_bv^{bk}{}_a \times \epsilon_{c_1c_2(k} \text{tr} (\bar{\phi}^{c_1} \bar{\phi}^{c_2} \bar{\phi}^d \psi_{d+}) \\ & + 36\epsilon_{a_1a_2(i} u^{a_1k} v^{a_2}{}_j) \times \left[2\text{tr} \left(\phi^{(i} \phi^c \phi^{j)} \psi_{(c} \psi_{k)} \right) + 2\text{tr} \left(\phi^{(i|} \phi^c \phi^{j|)} \psi_{(c} \psi_{k)} \right) \right. \\ & \quad \left. + 9\text{tr} \left(\phi^{(i} \phi^j \psi_{(c} \phi^c \psi_{k)} \right) - 6\text{tr} \left(\phi^{(i} \phi^{j)} \psi_{(c} \phi^c \psi_{k)} \right) \right] \\ & - 9\epsilon_{a_1a_2j} u^{a_1b} v^{a_2}{}_b \times \left[2\text{tr} \left(\phi^{(j} \phi^c \phi^{d)} \psi_{(c} \psi_{d)} \right) + 2\text{tr} \left(\phi^{(j|} \phi^c \phi^{d|)} \psi_{(c} \psi_{d)} \right) \right. \\ & \quad \left. + 9\text{tr} \left(\phi^{(j} \phi^d \psi_{(c} \phi^c \psi_{d)} \right) - 6\text{tr} \left(\phi^{(j} \phi^{d)} \psi_{(c} \phi^c \psi_{d)} \right) \right] \\ & - 20u^{a(i} v^{j)}{}_a \times \epsilon_{a_1a_2a_3} \left[2\text{tr} (\psi_{(i} \psi_{j)} \phi^{a_1} \phi^{a_2} \phi^{a_3}) + \text{tr} (\psi_{(i} \phi^{a_1} \psi_{j)} \phi^{a_2} \phi^{a_3}) \right] \\ & - 36u^{a(i} v^{j)}{}_a \times \epsilon_{a_1a_2(i} \left[\text{tr} (\psi_{j)} \psi_{a_3} \phi^{a_1} \phi^{a_2} \phi^{a_3}) + \text{tr} (\psi_{j)} \psi_{a_3} \phi^{a_1} \phi^{a_3} \phi^{a_2}) + \text{tr} (\psi_{j)} \psi_{a_3} \phi^{a_3} \phi^{a_1} \phi^{a_2}) \right] \\ & - 36u^{a(i} v^{j)}{}_a \times \epsilon_{a_1a_2(i} \left[\text{tr} (\psi_{j)} \phi^{a_1} \psi_{a_3} \phi^{a_2} \phi^{a_3}) + \text{tr} (\psi_{j)} \phi^{a_1} \psi_{a_3} \phi^{a_3} \phi^{a_2}) + \text{tr} (\psi_{j)} \phi^{a_3} \psi_{a_3} \phi^{a_1} \phi^{a_2}) \right] \\ & - 36u^{a(i} v^{j)}{}_a \times \epsilon_{a_1a_2(i} \left[\text{tr} (\psi_{j)} \phi^{a_1} \phi^{a_2} \psi_{a_3} \phi^{a_3}) + \text{tr} (\psi_{j)} \phi^{a_1} \phi^{a_3} \psi_{a_3} \phi^{a_2}) + \text{tr} (\psi_{j)} \phi^{a_3} \phi^{a_1} \psi_{a_3} \phi^{a_2}) \right] \\ & - 36u^{a(i} v^{j)}{}_a \times \epsilon_{a_1a_2(i} \left[\text{tr} (\psi_{j)} \phi^{a_1} \phi^{a_2} \phi^{a_3} \psi_{a_3}) + \text{tr} (\psi_{j)} \phi^{a_1} \phi^{a_3} \phi^{a_2} \psi_{a_3}) + \text{tr} (\psi_{j)} \phi^{a_3} \phi^{a_1} \phi^{a_2} \psi_{a_3}) \right] \\ & + 60u^{a(i} v^{j)}{}_a \times \epsilon_{a_1a_2(i} \left[\text{tr} (\psi_{j)} \phi^{a_1} \phi^{a_2}) \text{tr} (\psi_{a_3} \phi^{a_3}) + \text{tr} (\psi_{j)} \phi^{a_1} \phi^{a_3}) \text{tr} (\psi_{a_3} \phi^{a_2}) + \text{tr} (\psi_{j)} \phi^{a_3} \phi^{a_1}) \text{tr} (\psi_{a_3} \phi^{a_2}) \right] \end{aligned}$$

$$E = \frac{21}{2} , \quad R_1 = R_2 = R_3 = \frac{5}{2} , \quad J_1 = J_2 = \frac{3}{2}$$

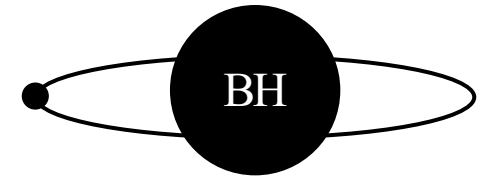
and infinitely many other examples & WIP in SU(4) SYM

Non-graviton cohomology

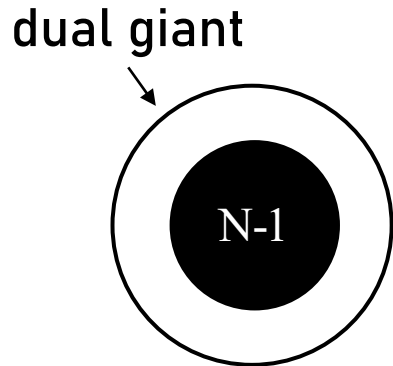
In $SU(3)$ SYM, we also constructed many other cohomologies

- graviton hair: infinitely many are allowed hair

$$O_{\text{BH}} \cdot \partial_{++}^{n+} \partial_{+-}^{n-} \text{tr}(\bar{\phi}^2)$$



- dual giant graviton hair



dual giant graviton reduce RR flux inside

→ $SU(N-1)$ core BH with dual giant hair

$$(O_3)^{i_1 \cdots i_n} \xrightarrow{SU(2) \times U(1)} \hat{\phi}^{i_1 \cdots i_n} O_2 + \cdots$$

↑
SU(2) threshold

Conclusion

We studied BPS spectrum of 4d $N=4$ $SU(N)$ SYM

- constructed non-graviton operators for $N=3$ at weak coupling

Now we can view black holes as an ensemble of microstates

- computing 2pt functions in the black hole background ~ information paradox

[Ahn, Kim, Lee, Lee, Yoon WIP]

- microscopically understanding the black hole hairs

Increasing N & higher-loop correction

- recently, it has been shown that **quantum correction** to supercharge can lift the black hole cohomology

[Choi, Lee '25] [Kim, Kim, Lee, Park '26]

- it is open question which cohomology survive at higher-loop correction

Thank you

Appendix

Maximal Super-Yang-Mills

The $\mathcal{N} = 4$ super-Yang-Mills theory with $SU(N)$ gauge group on R^4

- 3 chiral multiplets $\phi_m, \bar{\phi}^m, \psi_{m\alpha}, \bar{\psi}^m_{\dot{\alpha}}$ ($m = 1, 2, 3$)
- 1 vector multiplet $A_\mu \sim A_{\alpha\dot{\beta}}, \lambda_\alpha, \bar{\lambda}_{\dot{\alpha}}$ ($\alpha, \dot{\alpha} = \pm$)

Operators in R^4
 $\updownarrow$
 States in $S^3 \times R$

Among 32 supercharges $Q^I_\alpha, \bar{Q}_{I\dot{\alpha}}, S^\alpha_I, \bar{S}^{I\dot{\alpha}}$ ($I = 1, 2, 3, 4$)

we choose specific $Q \equiv Q^4_-$, $S \equiv S^-_4$ which BPS operator preserves
 (the same SUSY dual black hole preserves)

Such operators are called $\frac{1}{16}$ -BPS operators ($\frac{1}{16} = \frac{2}{32}$)

$$[Q, O] = 0 \ \& \ [S, O] = 0 \quad \Leftrightarrow \quad [\{Q, S\}, O] = 0$$

- saturate the BPS bound $E - (R_1 + R_2 + R_3 + J_1 + J_2) = 0$

which follows from the algebra $2\{Q, S\} = E - (R_1 + R_2 + R_3 + J_1 + J_2) \geq 0$

Graviton cohomology

There is a well-known class of cohomology called **graviton** [Gunaydin, Marcus '84]
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which corresponds to BPS Kaluza-Klein gravitons in $\text{AdS}_5 \times S^5$

- Chiral primary $\text{tr}(\bar{\phi}^{(i_1} \dots \bar{\phi}^{i_n)})$

Since $Q\bar{\phi}^m = 0$, **Q -closed**.

Due to symmetrization, it is **not Q -exact**. Recall $Q\psi_{m+} = -ig_{YM}\epsilon_{mnp}[\bar{\phi}^n, \bar{\phi}^p]$

- Superconformal descendants

By acting supercharges $Q_+^m, \bar{Q}_{m\dot{\alpha}}$ and BPS derivatives $\partial_{+\dot{\alpha}}$ on chiral primary

- Multiplying the above single trace gravitons
- Q -closed for any N

Black hole cohomology

In interacting supersymmetric theories, trace relations can **generate new operators**

$$Q\mathcal{O} = (\text{Trace relation}) = 0$$

- explains the rest of the cohomology, which we call non-graviton cohomology

These are possible candidate for black hole microstates

- since microstates of black holes are heavy $E \sim 1/G_N \sim N^2$ & deconfined d.o.f.
 - ← subjected to trace relations
- we often call them black hole cohomology

This phenomena is also known as fortuity [Chang, Lin '24]

Conclusion

We studied BPS spectrum of 4d $N=4$ $SU(N)$ SYM

- Computed index of non-gravitons for $N=3,4$ at weak coupling
- Constructed black hole operators for $N=3$

So far, we discussed classical cohomology

- recently, it has been shown that **quantum correction** to supercharge can lift the black hole cohomology [Choi-Lee '25]
- it is open question which cohomology survive at higher-loop correction

Cohomology problem can be studied in various theories

- SYK model [Chang-Chen-Sia-Yang '24] D1-D5 CFT [Chang-Lin-Zhang '25]
ABJM [Belin-Singh-Vadala-Zaffaroni '25] [Behan-de Gioia '25] ABJ [Kim-JL-Lee-Oh '25] and LS theory → See Seunggyu's talk