

Effective Field Theory of coupled dark energy and dark matter

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Motivation and outline

⇒ Λ CDM is a successful model

- **Cosmological constant problem**: why the value of CC is very small.
- **Coincidence problem**: why DE dominates at very recent redshift?
- **Observational Anomalies**: Hubble tension, S_8 anomaly, and DESI observation.

⇒ Dynamical DE like, Quintessence, k-essence, proca field were introduced to address them.

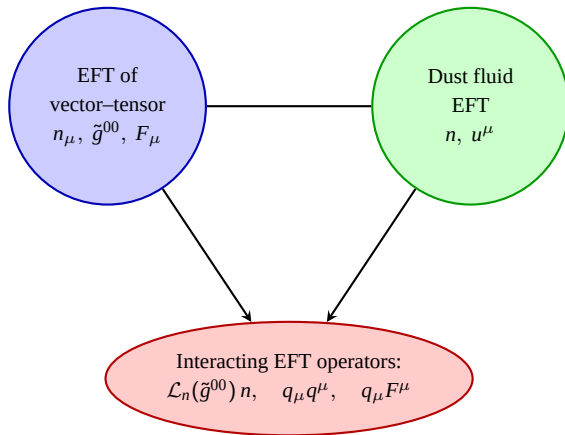
⇒ **Interacting dark energy and dark matter** is introduced to address the coincidence problem.

⇒ Interacting dark energy and dark matter is shown to be successful to address **H_0 , and S_8 tension**, as well as the **DESI observations**.

⇒ There are many models of interacting DE and DM, some of which predict a **suppressed G_{eff}** , potentially reducing structure growth and **addressing tension** in large-scale observations.

Motivation and outline

- ⇒ Can we formulate a **theory agnostic** interacting description of Dark sector?
- ⇒ The phenomenology can be approached either via **phenomenological parameterizations** or by constructing a **unified EFT framework**, which captures all allowed operators consistent with symmetry principles.
- ⇒ Our EFT approach yields a **well-defined, stable, and predictive theory**, providing explicit stability conditions and controlled perturbation evolution, enabling meaningful inference from the data.



EFT of vector tensor: Symmetry breaking pattern

We assume existence of **preferred direction** by timelike vector, v^μ , which spontaneously break spacetime symmetry

$$v_\mu = \partial_\mu \tilde{t} + g_M A_\mu$$

$\tilde{t}$ is Stückelberg field associated with the $U(1)$ transformation of the vector field A_μ , and g_M is the gauge coupling.

Defining $\tilde{t} = t$ the time coordinate, the preferred vector is

$$v_\mu = \tilde{\delta}_\mu^0 \equiv \delta_\mu^0 + g_M A_\mu .$$

The **residual symmetries** are

$$\begin{aligned} \text{Combinet time and } U(1) & : t \rightarrow t - g_M \chi , \quad A_\mu \rightarrow A_\mu + \partial_\mu \chi , \\ \text{Space(diff)} & : \vec{x} \rightarrow \vec{x}'(t, \vec{x}) \end{aligned}$$

preferred vector $\neq$ preferred slicing

EFT of vector tensor: Building blocks

Norm of the preferred vector

$$\tilde{g}^{00} = \tilde{\delta}_\mu^0 \tilde{\delta}_\nu^0 g^{\mu\nu}$$

We can define a **time like vector**

$$\tilde{n}_\mu \equiv \frac{\tilde{\delta}_\mu^0}{\sqrt{-\tilde{g}^{00}}}, \quad \tilde{n}_\alpha \tilde{n}^\alpha = -1$$

Then a **projection tensor** can be defined

$$\tilde{h}_{\mu\nu} = g_{\mu\nu} + \tilde{n}_\mu \tilde{n}_\nu.$$

$\tilde{n}_\mu$ is not hypersurface orthogonal

The derivative of the preferred vector can have two parts

$$\nabla_\mu \tilde{\delta}_\nu^0 = \frac{\tilde{n}_\nu \nabla_\mu \tilde{g}^{00}}{2\sqrt{-\tilde{g}^{00}}} - \sqrt{-\tilde{g}^{00}} \nabla_\mu \tilde{n}_\nu.$$

$\nabla_\mu \tilde{n}_\nu \Rightarrow$ decomposed in to

$$\tilde{K}_{\mu\nu} = \tilde{h}_{(\mu}{}^\alpha \nabla_\alpha \tilde{n}_{\nu)},$$

$$\tilde{\omega}_{\mu\nu} = \tilde{h}_{[\mu}{}^\alpha \nabla_\alpha \tilde{n}_{\nu]},$$

$$\tilde{a}_\mu = \tilde{n}^\alpha \nabla_\alpha \tilde{n}_\mu.$$

$\tilde{K}_{\mu\nu}$, $\tilde{\omega}_{\mu\nu}$, $\tilde{a}_\mu$, specify the geometric properties of the preferred vector.

EFT of vector tensor: Building blocks

For the gauge field A_μ , we have the **field strength**

$$F_{\mu\nu} = \nabla_\mu A_\nu - \nabla_\nu A_\mu$$

Decomposed in to

$$\begin{aligned} \text{Electric} &: \quad \tilde{F}_\mu \equiv \tilde{h}^\alpha F_{\mu\alpha}, \\ \text{Magnetic} &: \quad \tilde{F}_{\mu\nu} \equiv \tilde{h}_\mu^\alpha \tilde{h}_\nu^\beta F_{\alpha\beta}. \end{aligned}$$

We have the follow relation

$$\begin{aligned} \tilde{\omega}_{\mu\nu} &= -\frac{g_M}{2\sqrt{-\tilde{g}^{00}}} \tilde{F}_{\mu\nu}, \\ \tilde{a}_\mu &= \frac{g_M}{\sqrt{-\tilde{g}^{00}}} \tilde{F}_\mu - \frac{\tilde{h}^\alpha_\mu \partial_\alpha \tilde{g}^{00}}{2\tilde{g}^{00}}. \end{aligned}$$

Hence $\tilde{\omega}_{\mu\nu}$, and a_μ are redundant.

We can have **derivative operator**

$$\tilde{D}_\mu T^{\nu\dots}_{\rho\dots} = \tilde{h}^\alpha_\mu \tilde{h}^\nu_\beta \dots \tilde{h}^\gamma_\rho \nabla_\alpha T^{\beta\dots}_{\gamma\dots}.$$

Then the **gauss equation** can be written

$$\begin{aligned} \tilde{h}^\alpha_\mu \tilde{h}^\beta_\nu \tilde{h}^\gamma_\rho \tilde{h}^\delta_\sigma R_{\alpha\beta\gamma\delta} &= {}^{(3)}\tilde{R}_{\mu\nu\rho\sigma} + 2\tilde{B}_{[\rho|\mu}\tilde{B}_{|\sigma]\nu}, \\ \tilde{B}_{\mu\nu} &= \tilde{K}_{\mu\nu} + \tilde{\omega}_{\mu\nu}. \end{aligned}$$

Hence

$${}^{(3)}\tilde{R}_{\mu\nu} = {}^{(3)}\tilde{R}^\alpha_{\mu\alpha\nu}, \quad {}^{(3)}\tilde{R} = {}^{(3)}\tilde{R}^\mu_\mu = {}^{(3)}\tilde{R}^{\mu\nu}_{\mu\nu},$$

and

$$R = {}^{(3)}\tilde{R} + \tilde{B}_{\mu\nu}\tilde{B}^{\mu\nu} - \tilde{K}^2 - 2\nabla_\mu(\tilde{a}^\mu - \tilde{K}\tilde{n}^\mu).$$

EFT of vector tensor: Building blocks

The EFT building blocks are

$$\tilde{g}^{00}, \quad {}^{(3)}\tilde{R}_{\mu\nu\rho\sigma}, \quad \tilde{F}_{\mu\nu}, \quad \tilde{F}_\mu, \quad \tilde{K}_{\mu\nu}, \\ \mathfrak{L}_{\tilde{n}}, \quad \tilde{D}_\mu, \quad \text{and} \quad g^{\mu\nu} \text{ (or } h^{\mu\nu})$$

The general action

The most general action in **unitary gauge** is

$$S = \int d^4x \mathcal{L}_{\text{DE}} + S_{\text{m}}[\psi, g]$$

where

$$\mathcal{L}_{\text{DE}} = \mathcal{L}_{\text{DE}}(\tilde{g}^{00}, {}^{(3)}\tilde{R}_{\alpha\beta\gamma\delta}, \tilde{F}_{\mu\nu}, \tilde{F}_\mu, \tilde{K}_{\mu\nu}, \mathfrak{L}_{\tilde{n}}, \tilde{D}_\mu)$$

With no transverse degrees of freedom

Without transverse mode, gauge field is

$$A_\mu = [A_0, \partial_i A]$$

We can use the residual gauge freedom of the combined $U(1)$ and time diff to set $A = 0$.

Then

$$A_\mu = [A_0, \mathbf{0}].$$

In this case the temporal freedom is completely fixed except the time reparameterization $t \rightarrow t'(t)$.

The vector $\tilde{n}_\mu$ coincide with the vector orthogonal to the hypersurface, n_μ .

Then

$$\tilde{n}^\mu = n^\mu, \tilde{D}_\mu = D_\mu, \tilde{K}_{\mu\nu} = K_{\mu\nu}, {}^{(3)}\tilde{R}_{\alpha\beta\gamma\delta} = {}^{(3)}R_{\alpha\beta\gamma\delta}.$$

The preferred vector

$$v_\mu = \tilde{\delta}_\mu^0 = \delta_\mu^0 + g_M A_\mu = (1 + g_M A_0, \mathbf{0}).$$

The norm

$$\tilde{g}^{00} = \tilde{\delta}_\alpha^0 \tilde{\delta}_\beta^0 g^{\alpha\beta} = (1 + g_M A_0)^2 g^{00}.$$

Hence, the building blocks becomes

$$n_\mu, \tilde{g}^{00}, F_\mu, K_{\mu\nu}, {}^{(3)}R_{\mu\nu}.$$

Luminal EFT of Dark energy action

Using the above ansatz and considering the **Luminal EFT of vector-tensor DE**: $c_T = c$, where c speed of light

$$\begin{aligned}\mathcal{L}_{\text{DE}} = & \frac{M_*^2}{2} f(t) \left[{}^{(3)}R + K_{\mu\nu} K^{\mu\nu} - K^2 \right] - \hat{\Lambda}(t) - \hat{c}(t) \tilde{g}^{00} - d(t) K \\ & + \frac{1}{2} \hat{M}_2^4(t) \left(\frac{\delta \tilde{g}^{00}}{-\tilde{g}_{\text{BG}}^{00}} \right)^2 - \frac{1}{2} \bar{M}_1^3(t) \left(\frac{\delta \tilde{g}^{00}}{-\tilde{g}_{\text{BG}}^{00}} \right) \delta K + \frac{1}{2} \gamma_1(t) F_\mu F^\mu ,\end{aligned}$$

In this case we have the following **consistency conditions**

$$\begin{aligned}\dot{\hat{\Lambda}} + 3H\dot{d} + \dot{\hat{c}}\tilde{g}_{\text{BG}}^{00} &= 0 , \\ 2\hat{M}_2^4 \frac{d}{\bar{N}dt} \ln(-\tilde{g}_{\text{BG}}^{00}) + 3\bar{M}_1^3 \dot{H} + 2\dot{\hat{c}}\tilde{g}_{\text{BG}}^{00} &= 0 , \\ \dot{d} + \frac{1}{2} \bar{M}_1^3 \frac{d}{\bar{N}dt} \ln(-\tilde{g}_{\text{BG}}^{00}) &= 0 , \\ \dot{f} &= 0 .\end{aligned}$$

Dark Matter action

Here we assume that that dust fluid is described by three scalars

$$\phi^i(t, \mathbf{x}), \quad i = 1, 2, 3$$

these scalar fields can be understood as comoving coordinates $\phi^i = x^i$.

Fluid is described by the invariance under **volume preserving diffs**

$$\phi^i \rightarrow \phi'^i \quad \text{s.t.} \quad \det \frac{\partial \phi'^i}{\partial \phi^i} = 1.$$

$$\mathcal{S}_{\text{DM}} = \int d^4x \sqrt{-g} \mathcal{L}_{\text{DM}}, \quad \mathcal{L}_{\text{DM}} = -\hat{m}_c n, \quad \hat{m}_c \text{ is a constant.}$$

the invariant building blocks are

$$n \equiv \sqrt{\det g^{\mu\nu} \partial_\mu \phi^i \partial_\nu \phi^i}$$
$$u^\mu \equiv -\frac{1}{6} \epsilon_{ijk} \epsilon^{\mu\nu\rho\sigma} \partial_\nu \phi^i \partial_\rho \phi^j \partial_\sigma \phi^k.$$

In this formulation the current $\mathcal{J}^\mu \equiv n u^\mu$ **conservation is off-shell**

$$\nabla_\mu \mathcal{J}^\mu = 0$$

in the comoving gauge $\phi^i = x^i$

$$n = \sqrt{\det g^{ij}}, \quad u^\mu = \frac{\delta^\mu_0}{\sqrt{-g_{00}}}$$

Interacting action

Here are the basic ingredients from dark sectors

$$\underbrace{n_\mu, \tilde{g}^{00}, F_\mu}_{\text{DE}}, \quad \underbrace{n, u^\mu}_{\text{DM}}.$$

The possible interactions are

$$\mathcal{L}_n(\tilde{g}^{00})n, \quad n_\mu u^\mu, \quad F_\mu u^\mu.$$

The first one is energy transfer, the second two are momentum transfer.

Introducing $\mathcal{U} \equiv n_\mu u^\mu$, and $q^\mu \equiv u^\mu + n^\mu \mathcal{U}$.

Related by

$$q_\mu q^\mu = -1 + \mathcal{U}^2$$

In comoving gauge

$$q^\mu = \left(0, \frac{N^i}{\sqrt{N^2 - N_i N^i}} \right).$$

We use q_μ instead of u_μ , hence interactions are

$$q_\mu q^\mu, \quad q_\mu F^\mu.$$

Interacting action

Then the general interacting Lagrangian is given by

$$\begin{aligned}\mathcal{L}_{\text{int}} &= \mathcal{L}_n(\tilde{g}^{00})n + \mathcal{L}_{q^2}(\tilde{g}^{00})q^\mu q_\mu + \mathcal{L}_{q \cdot F}(\tilde{g}^{00})F_\mu q^\mu + \cdots \\ &= -\Delta\Lambda(t) - \Delta c(t)\tilde{g}^{00} - \Delta m_c(t)n \\ &\quad + \frac{1}{2}\Delta M_2^4(t) \left(\frac{\delta\tilde{g}^{00}}{-\tilde{g}_{\text{BG}}^{00}} \right)^2 - m_1^4(t) \frac{\delta n}{\bar{n}} \left(\frac{\delta\tilde{g}^{00}}{-\tilde{g}_{\text{BG}}^{00}} \right) - m_2^4(t)q^\mu q_\mu - \bar{m}_1^2(t)q^\mu F_\mu \\ &\quad + \cdots ,\end{aligned}$$

where

$$\begin{aligned}\Delta\Lambda(t) &= \bar{\mathcal{L}}_{n\tilde{g}^{00}} \bar{n} \tilde{g}_{\text{BG}}^{00} , & \Delta c(t) &= -\bar{\mathcal{L}}_{n\tilde{g}^{00}} \bar{n} , & \Delta m_c(t) &= -\bar{\mathcal{L}}_n , \\ \Delta M_2^4(t) &= \bar{\mathcal{L}}_{n\tilde{g}^{00}\tilde{g}^{00}} \bar{n} (\tilde{g}_{\text{BG}}^{00})^2 , & m_1^4(t) &= \bar{\mathcal{L}}_{n\tilde{g}^{00}} \tilde{g}_{\text{BG}}^{00} \bar{n} , \\ m_2^4(t) &= -\bar{\mathcal{L}}_{q^2} , & \bar{m}_1^2(t) &= -\bar{\mathcal{L}}_{q \cdot F} ,\end{aligned}$$

and

$$\delta n = n - \bar{n}(t) , \quad \mathcal{L}_{n\tilde{g}^{00}} = \frac{d\mathcal{L}_n}{d\tilde{g}^{00}} , \quad \mathcal{L}_{n\tilde{g}^{00}\tilde{g}^{00}} = \frac{d\mathcal{L}_{n\tilde{g}^{00}}}{d\tilde{g}^{00}} .$$

EFT action for coupled Dark Energy and Dark Matter

Action

$$\mathcal{S} = \int d^4x \sqrt{-g} \left(\mathcal{L}_D^{\text{NL}} + \mathcal{L}_D^{(2)} \right) + \mathcal{S}_m ,$$

$$\mathcal{L}_D^{\text{NL}} = \frac{M_*^2}{2} f(t) \left({}^{(3)}R + K_{\mu\nu} K^{\mu\nu} - K^2 \right) - \Lambda(t) - \tilde{c}(t) \tilde{g}^{00} - d(t) K - m_c(t) n ,$$

$$\mathcal{L}_D^{(2)} = \frac{1}{2} M_2^4(t) \left(\frac{\delta \tilde{g}^{00}}{-\tilde{g}_{\text{BG}}^{00}} \right)^2 - \frac{1}{2} \bar{M}_1^3(t) \left(\frac{\delta \tilde{g}^{00}}{-\tilde{g}_{\text{BG}}^{00}} \right) \delta K + \frac{1}{2} \gamma_1(t) F_\mu F^\mu$$

$$- m_1^4(t) \frac{\delta n}{\bar{n}} \left(\frac{\delta \tilde{g}^{00}}{-\tilde{g}_{\text{BG}}^{00}} \right) - m_2^4(t) q^\mu q_\mu - \bar{m}_1^2(t) q^\mu F_\mu .$$

$$\dot{\rho}_c + 3H\rho_c = \frac{d}{dt} \ln(m_c) \bar{\rho}_c , \quad \rho_c \equiv m_c \bar{n}_c .$$

For the Horndeski case

$$\mathcal{L}_{\text{DM}} + \mathcal{L}_{\text{int}} = -f_1(t, X, Z)n + f_2(t, X, Z) .$$

where $X = -\frac{1}{2}\nabla_\mu\phi\nabla^\mu\phi$, and $Z = -u^\mu\nabla_\mu\phi$

The mapping for interacting coefficient

$$\begin{aligned} m_c &= f_1 , \\ m_1^4 &= -\bar{n} \left(f_{1,X} X_{\text{BG}} + \frac{1}{2} \sqrt{2X_{\text{BG}}} f_{1,Z} \right) , \\ m_2^4 &= \frac{1}{2} \sqrt{2X_{\text{BG}}} (f_{1,Z} \bar{n} - f_{2,Z}) , \\ g_M &= \gamma_1 = \bar{m}_1^2 = 0 . \end{aligned}$$

For GP theory

$$\mathcal{L}_{\text{DM}} + \mathcal{L}_{\text{int}} = -f_1(\tilde{X}, \tilde{Z}, \tilde{E})n + f_2(\tilde{X}, \tilde{Z}, \tilde{E}) ,$$

$\tilde{X} = -A_\mu A^\mu / 2$, $\tilde{Z} = A_\mu u^\mu$, and $\tilde{E} = -A^\mu F_{\mu\nu} u^\nu$.

The mapping

$$\begin{aligned} \gamma_1 &= 1 , \\ m_c &= f_1 , \\ m_1^4 &= -\bar{n} \left(f_{1,\tilde{X}} \tilde{X}_{\text{BG}} + \frac{1}{2} \sqrt{2\tilde{X}_{\text{BG}}} f_{1,\tilde{Z}} \right) , \\ m_2^4 &= \frac{1}{2} \sqrt{2\tilde{X}_{\text{BG}}} (f_{1,\tilde{Z}} \bar{n} - f_{2,\tilde{Z}}) , \\ \bar{m}_1^2 &= -\sqrt{2\tilde{X}_{\text{BG}}} (f_{1,\tilde{E}} \bar{n} - f_{2,\tilde{E}}) . \end{aligned}$$

[A. De Felice, S. Nakamura S. Tsujikawa- arXiv:2004.09384, MCP, K. Koyama-arXiv:2405.06565]

EFT action in α - Basis

From the definition

$$\frac{\delta \tilde{g}^{00}}{-\tilde{g}_{\text{BG}}^{00}} = \frac{\delta g^{00}}{-g_{\text{BG}}^{00}} - \frac{2g_M}{(1 + g_M A_0)} \delta A_0, \quad F_\mu F^\mu = (n^0 \nabla_i A_0)(n^0 \nabla^i A_0) = \frac{\delta^{ij} \nabla_i \delta A_0 \nabla_j \delta A_0}{a^2 \bar{N}^2},$$

$$q^\mu q_\mu = \frac{N^i N_i}{N^2 - N_j N^j} \simeq \frac{N^i N_i}{\bar{N}^2}, \quad q^\mu F_\mu = \frac{n^0 N^i \nabla_i \delta A_0}{\sqrt{N^2 - N_j N^j}} \simeq \frac{N^i \nabla_i \delta A_0}{\bar{N}^2}.$$

We can write the Lagrangian as

$$\mathcal{L}_{\text{D}}^{(2)} = \mathcal{L}_{\text{ST}}^{(2)} + \mathcal{L}_{\delta \hat{A}_0}^{(2)},$$

where

$$\begin{aligned} \mathcal{L}_{\text{ST}}^{(2)} &= 2M_2^4 \left(\frac{\delta N}{\bar{N}} \right)^2 - \bar{M}_1^3 \frac{\delta N}{\bar{N}} \delta K - 2m_1^4 \frac{\delta N}{\bar{N}} \frac{\delta n}{\bar{n}} - m_2^4 \frac{N^i N_i}{\bar{N}^2}, \\ \mathcal{L}_{\delta \hat{A}_0}^{(2)} &= \frac{1}{2} (\nabla_i \delta \hat{A}_0 \nabla^i \delta \hat{A}_0 + g_{\text{eff}}^2 M_2^4 \delta \hat{A}_0^2) \\ &\quad + \left[\frac{\bar{m}_1^2 \nabla_i N^i}{\sqrt{\gamma_1} \bar{N}} + g_{\text{eff}} \left(\frac{1}{2} \bar{M}_1^3 \delta K - 2M_2^4 \frac{\delta N}{\bar{N}} + m_1^4 \frac{\delta n}{\bar{n}} \right) \right] \delta \hat{A}_0, \quad g_{\text{eff}} \equiv \frac{2g_M \bar{N}}{\sqrt{\gamma_1} (1 + g_M \bar{A}_0)}, \quad \delta \hat{A}_0 \equiv \frac{\sqrt{\gamma_1}}{\bar{N}} \delta A_0. \end{aligned}$$

EFT action in α -Basis

After **integrating out** $\delta\hat{A}_0$, we can write the EFT action in real space,

$$\begin{aligned}
 \mathcal{S}_D^{(2)} &= \int d^4x \sqrt{-g} \mathcal{L}_D^{(2)} \\
 &= \int d^4x \bar{N} a^3 \left[2M_{2,\text{eff}}^4 \left(\frac{\delta N}{\bar{N}} \right)^2 - \bar{M}_{1,\text{eff}}^3 \frac{\delta N}{\bar{N}} \delta K - 2m_{1,\text{eff}}^4 \frac{\delta N}{\bar{N}} \frac{\delta n}{\bar{n}} - m_2^4 \frac{N_i N^i}{\bar{N}^2} \right. \\
 &\quad \left. - \bar{m}_{1,\text{eff}}^4 \frac{(\nabla_i N^i)^2}{\bar{N}^2} - \frac{\mu_1^4}{2} \left(\frac{\delta n}{\bar{n}} \right)^2 - \frac{\mu_2^2}{8} (\delta K)^2 - \frac{\mu_3^3}{2} \delta K \frac{\delta n}{\bar{n}} - \frac{\mu_4^4}{2} \delta K \frac{\nabla_i N^i}{\bar{N}} \right. \\
 &\quad \left. + 2\mu_5^5 \frac{\delta N}{\bar{N}} \frac{\nabla_i N^i}{\bar{N}} - \mu_6^5 \frac{\delta n}{\bar{n}} \frac{\nabla_i N^i}{\bar{N}} \right],
 \end{aligned}$$

where

$$\begin{aligned}
 M_{2,\text{eff}}^4 &\equiv (1 - \mathcal{G}) M_2^4, & \bar{M}_{1,\text{eff}}^3 &\equiv (1 - \mathcal{G}) \bar{M}_1^3, & m_{1,\text{eff}}^4 &\equiv (1 - \mathcal{G}) m_1^4, \\
 \bar{m}_{1,\text{eff}}^4 &\equiv \frac{\mathcal{G} \bar{m}_1^4}{2\gamma_1 g_{\text{eff}}^2 M_2^4}, & \mu_1^4 &\equiv \mathcal{G} \frac{m_1^8}{M_2^4}, & \mu_2^2 &\equiv \mathcal{G} \frac{\bar{M}_1^6}{M_2^4}, & \mu_3^3 &\equiv \mathcal{G} \frac{\bar{M}_1^3 m_1^4}{M_2^4}, \\
 \mu_4^4 &\equiv \frac{\mathcal{G} \bar{M}_1^3 \bar{m}_1^2}{\sqrt{\gamma_1} g_{\text{eff}} M_2^4}, & \mu_5^5 &\equiv \frac{\mathcal{G} \bar{m}_1^2}{\sqrt{\gamma_1} g_{\text{eff}}}, & \mu_6^5 &\equiv \frac{\mathcal{G} m_1^4 \bar{m}_1^2}{\sqrt{\gamma_1} g_{\text{eff}} M_2^4}, & \mathcal{G} &\equiv \frac{g_{\text{eff}}^2 M_2^4}{-\nabla_i \nabla^i + g_{\text{eff}}^2 M_2^4}.
 \end{aligned}$$

As in the case of the EFT of uncoupled vector-tensor DE theories, we have three unified descriptions of the EFT of coupled DE and DM:

1. $g_{\text{eff}} \neq 0$, it corresponds to the **EFT of coupled vector DE**,
2. $g_{\text{eff}} = 0$, together with the consistency conditions, we obtain the **EFT of coupled scalar DE in shift-symmetric scalar-tensor theories**,
3. $g_{\text{eff}} = 0$, without the consistency conditions, we have the **EFT of coupled scalar DE in non-shift-symmetric scalar-tensor theories**.

Action in α - Basis

$$\begin{aligned}
 \mathcal{S}^{(2)} = & \int d^4x \bar{N} a^3 \frac{M^2}{2} \left[\delta K_{\mu\nu} \delta K^{\mu\nu} - \delta K^2 + \delta_2 \left(\frac{\sqrt{h}}{a^3} {}^{(3)}R \right) + \frac{\delta N}{\bar{N}} \delta^{(3)}R - 6\Omega_c H^2 \delta_2 \left(\frac{\sqrt{h}}{a^3} \frac{n}{\bar{n}} \right) \right. \\
 & + H^2 \tilde{\alpha}_K \left(\frac{\delta N}{\bar{N}} \right)^2 + 4\tilde{\alpha}_B H \delta K \frac{\delta N}{\bar{N}} + (\tilde{\alpha}_{m_1} - 6\Omega_c) H^2 \frac{\delta N}{\bar{N}} \frac{\delta n}{\bar{n}} + \alpha_{m_2} H^2 \frac{N^i N_i}{\bar{N}^2} \\
 & + \tilde{\alpha}_{\bar{m}_1} \frac{(\nabla_i N^i)^2}{\bar{N}^2} + \alpha_{\mu_1} H^2 \left(\frac{\delta n}{\bar{n}} \right)^2 + \alpha_{\mu_2} \delta K^2 + \alpha_{\mu_3} H \delta K \frac{\delta n}{\bar{n}} \\
 & \left. - \alpha_{\mu_4} \delta K \frac{\nabla_i N^i}{\bar{N}} - \alpha_{\mu_5} H \frac{\delta N}{\bar{N}} \frac{\nabla_i N^i}{\bar{N}} - \alpha_{\mu_6} H \frac{\delta n}{\bar{n}} \frac{\nabla_i N^i}{\bar{N}} \right] + \tilde{\mathcal{S}}_{\text{m}}^{(2)},
 \end{aligned}$$

$$\tilde{\mathcal{S}}_{\text{m}}^{(2)} \equiv \mathcal{S}_{\text{m}}^{(2)} + \delta_2 \int d^4x \bar{N} \sqrt{h} \left(\frac{\delta N}{\bar{N}} \bar{\rho}_{\text{m}} - \bar{p}_{\text{m}} \right),$$

Ghost conditions

$$N = \bar{N}(t), \quad N_i = 0, \quad h_{ij} = a^2(t)(\delta_{ij} + \tilde{h}_{ij}), \quad \tilde{h}_{11} = -\tilde{h}_{22} = h_+(t, z), \quad \tilde{h}_{12} = \tilde{h}_{21} = h_\times(t, z).$$

Tensor modes

The action reduces to

$$\mathcal{S}_T^{(2)} = \sum_{\lambda=+, \times} \int \frac{dtd^3k}{(2\pi)^3} \bar{N} a^3 \frac{M^2}{4} \left(|\dot{h}_\lambda|^2 - \frac{k^2}{a^2} |h_\lambda|^2 \right),$$

where we have changed the action into Fourier space, and k is the Fourier mode.

Expected as we have focused on theories with the luminal speed of gravitational waves.

$$\ddot{h}_\lambda + H(3 + \alpha_M) \dot{h}_\lambda + \frac{k^2}{a^2} h_\lambda = 0, \quad \alpha_M \equiv \frac{2\dot{M}}{HM}.$$

$$N = \bar{N}(t) (1 + \alpha) , \quad N_i = \bar{N}(t) (\partial_i \chi) , \quad h_{ij} = a^2(t) [(1 + 2\zeta)\delta_{ij} + 2\partial_i \partial_j E] .$$

Scalar modes

write the action in the form

$$\mathcal{S}^{(2)} = \int \frac{dt d^3k}{(2\pi)^3} \bar{N} a^3 \left(\dot{\vec{\chi}}_k^t \mathbf{K} \dot{\vec{\chi}}_{-k} - \vec{\chi}_k^t \tilde{\mathbf{G}} \vec{\chi}_{-k} - \frac{k}{a} \vec{\chi}_k^t \mathbf{B} \dot{\vec{\chi}}_{-k} \right) ,$$

where

$$\vec{\chi}_k^t = \left[\frac{\delta_c(\mathbf{k})}{k}, \zeta(\mathbf{k}), \frac{\delta_m(\mathbf{k})}{k} \right] , \quad \text{and} \quad \tilde{\mathbf{G}} = \frac{k^2}{a^2} \mathbf{G} + \mathbf{M} .$$

$\mathbf{K}$ and $\tilde{\mathbf{G}}$ are symmetric 3×3 matrices and $\mathbf{B}$ is anti-symmetric matrix.

High k limit

$$q_c \equiv \frac{K_{11}}{a^2} = \frac{M^2 H^2}{2} (3\Omega_c + \alpha_{m_2} - \alpha_{\tilde{m}_1}^2) > 0,$$

$$q_s \equiv K_{22} = \frac{M^2 (6\alpha_B^2 + \tilde{\alpha}_K)}{2(1 + \alpha_B)^2} > 0,$$

with the strong energy condition for the standard matter

$$\bar{\rho}_m(1 + \omega_m) > 0.$$

High k limit

$$c_s^2 = \hat{c}_s^2 + \Delta c_s^2,$$

where

$$\begin{aligned}\hat{c}_s^2 &= \frac{G_{22}}{K_{22}} = -\frac{1}{(6\alpha_B^2 + \tilde{\alpha}_K)} \left[3\Omega_c + 3\Omega_m(1 + w_m) + 2(1 + \alpha_B)(\alpha_B - \alpha_M - \epsilon_H) + \frac{2\dot{\alpha}_B}{H} \right. \\ &\quad \left. - 4\alpha_B^2\alpha_g^2 + \alpha_{m_2} - \alpha_{\bar{m}_1}^2 - 4\alpha_g\alpha_B\alpha_{\bar{m}_1} \right], \\ \Delta c_s^2 &= \frac{B_{12}^2}{K_{11}K_{22}} = \frac{(\alpha_{m_1} + 2\alpha_{m_2} - 2\alpha_{\bar{m}_1}^2 - 4\alpha_g\alpha_B\alpha_{\bar{m}_1})^2}{4(3\Omega_c + \alpha_{m_2} - \alpha_{\bar{m}_1}^2)(6\alpha_B^2 + \tilde{\alpha}_K)}.\end{aligned}$$

To avoid Laplacian instabilities

$$c_s^2 > 0.$$

Effective gravitational coupling

Small-scale CDM evolution:

$$\ddot{\delta}_c + \mathcal{C}\dot{\delta}_c - 4\pi G_{\text{eff}}\bar{\rho}_c\delta_c \simeq 0$$

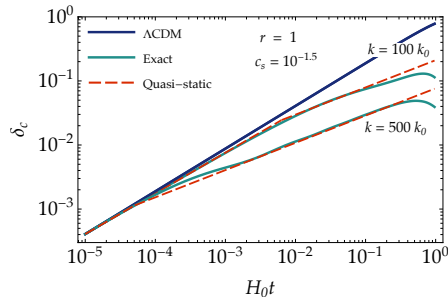
Damping coefficient:

$$\mathcal{C} = \frac{(5Hq_c + \dot{q}_c)\nu_s^2 + (2b_{12}\dot{b}_{12} + 5Hb_{12}^2)\nu_s - b_{12}^2\dot{\nu}_s}{q_c\nu_s^2 + b_{12}^2\nu_s}$$

Effective gravitational coupling:

$$G_{\text{eff}} = \frac{4\mu_{11}\nu_s^2 + 4g_{12}^2\nu_s - [3H^2(7 - 2\epsilon_H)b_{12}^2 + \dot{b}_{12}^2]\nu_s}{16\pi\bar{\rho}_c(q_c\nu_s^2 + b_{12}^2\nu_s)}$$

$$- \frac{2(\ddot{b}_{12} + 8H\dot{b}_{12} + 2\dot{g}_{12} + 4Hg_{12})b_{12}\nu_s - 2(\dot{b}_{12} + 3Hb_{12} + 2g_{12})b_{12}\dot{\nu}_s}{16\pi\bar{\rho}_c(q_c\nu_s^2 + b_{12}^2\nu_s)}$$



Summary & Future direction

Take-home message

- We successfully constructed **interacting EFT** of Dark Energy and Dark Matter
- This Int-EFT can accommodate Vector-Tensor Dark Energy, Scalar-Tensor Dark Energy (both shift symmetric and non-shift symmetric)
- We derived the **stability conditions** that theory should posses.
- We explored the **phenomenological impact** of the possible interactions.
- With some interaction terms, we can find a **possible suppression** for the growth of CDM ovedensity.

Future directions

EFT-param

Study the effect of interactions with different EFT α -parametrization.

Constraints

Parameter estimation with cosmological data, CMB and LSS.

Non-linear

Extending the formulation to non-linear scales.

Thank you!