Higgs inflation: with a Gauss-Bonnet term in Einstein Frame

@ Workshop on Cosmology and Quantum Space Time (CQUeST 2023) 임채호 교수님 추모 학회, July 31—August 04, 2023

Gansukh Tumurtushaa(JejuNU)

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arXiv:2308.00897

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coupling in the Jordan frame et term in the Einstein frame

Inflation

an idea of accelerated "exponential" expansion of the early universe driven by the so-called "inflaton" field ϕ





Inflation

an idea of accelerated "exponential" expansion of the early universe driven by the so-called "inflaton" field ϕ



Higgs Inflation identifies Standard Model "Higgs" field as the inflaton field ϕ .



Higgs Inflation with non-minimal coupling The action in the "Jordan" Frame:

$$S^{J} = \int d^{4}x \sqrt{-g^{J}} \left[\frac{M_{p}^{2}}{2} \Omega^{2}(\phi) R^{J} - \frac{1}{2} g_{ab}^{J} \nabla^{a} \phi \nabla^{b} \phi - \frac{1}{4} \lambda (\phi^{2} - \nu^{2})^{2} \right]$$

where the Higgs field is "non-minimally" coupled to gravity via

 $\Omega^2(\phi) =$

* σ non-minimal coupling constant

Phys. Lett. B 659 (2008), 703-706 Phys. Lett. B 675 (2009), 88-92 JHEP 01 (2011), 016 Phys. Rev. D 39 (1989), 399-404 Phys. Rev. D 41 (1990), 1783-1791 Phys. Rev. D 86 (2012)

$$= 1 + \frac{\sigma}{M_p^2} \phi^2.$$



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 $\Omega^2(\phi) =$

Einstein frame: the ϕ is "*minimally*" coupled to gravity R.

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$$= 1 + \frac{\sigma}{M_p^2} \phi^2.$$

One can move from the Jordan frame to the Einstein frame by using the conformal transformation (CT).





Higgs Inflation in the Einstein frame

Under the conformal transformation,

$$g^J_{ab} = \Omega^{-2} g_{ab}$$
, where $\Omega^2(\phi) = 1 + \frac{\sigma}{M_p^2} \phi^2$,

the "Einstein" frame action reads,

$$S = \int d^4x \sqrt{-g} \left[\frac{M_P^2}{2} R - \frac{1}{2} g^{ab} \nabla_a s \nabla_b s - \frac{\lambda M_P^4}{4\sigma^2} \left(1 - e^{-\sqrt{\frac{2}{3}} \frac{s}{M_P}} \right)^2 \right],$$

where $s/M_p \equiv \sqrt{3/2 \ln \Omega^2}$ is the canonically normalized scalar field.

Higgs Infl in the Einstein

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where $s/M_p \equiv \sqrt{3/2 \ln \Omega^2}$ is the canonically normalized scalar field.



Higgs Inflation in the Einstein frame

To be consistent with the CMB observation: $\lambda/\sigma^2 \simeq 5 \times 10^{-10}$



$$V(s) = \frac{\lambda M_p^4}{4\sigma^2} \left(1 - e^{-\sqrt{\frac{2}{3}}}\right)$$

Phys. Lett. B 659 (2008) 703, Phys. Rev. D 92 (2015) no.8, 083512 JHEP 02 (2021) 109

• indicating $\sigma \gg 1$ (unless the λ is tiny) \implies causes a strong coupling problem, where the model loses perturbative unitarity at a scale M_p/σ , well below the Planck scale.







While it is the most minimalistic and experimentally compatible, it is feasible to expect additional interactions to be present.

$$S^J = \int d^4x \sqrt{-g^J} \left[\frac{M_p^2}{2} \Omega^2(\phi) R^J \right]$$

Phys. Rev. D 93 (2016) no.6, 063519 Phys. Rev. D 107 (2023) no.6, 06350

 $-\frac{1}{2}g^J_{ab}\nabla^a\phi\nabla^b\phi - V(\phi) + \omega(\phi)R^{2^J}_{GB} \quad ,$



$$S^{J} = \int d^{4}x \sqrt{-g^{J}} \left[\frac{M_{p}^{2}}{2} \Omega^{2}(\phi) R^{J} - \frac{1}{2} g_{ab}^{J} \nabla^{a} \phi \nabla^{b} \phi - V(\phi) + \omega(\phi) R_{GB}^{2^{J}} \right]$$

where $R_{GB}^2 = R^2 - 4R_{\mu\nu}R^{\mu\nu} + R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma}$ is the Gauss-Bonnet combination and $\omega(\phi)$ is the coupling function, while

$$\Omega^{2}(\phi) = 1 + \frac{\sigma}{M_{p}^{2}}\phi^{2}, \quad V(\phi) = \frac{\lambda}{4}(\phi^{2} - \nu^{2})^{2}.$$

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$$S^{J} = \int d^{4}x \sqrt{-g^{J}} \left[\frac{M_{p}^{2}}{2} \Omega^{2}(\phi) R^{J} - \frac{1}{2} g_{ab}^{J} \nabla^{a} \phi \nabla^{b} \phi - V(\phi) + \omega(\phi) R_{GB}^{2^{J}} \right]$$

where $R_{GR}^2 = R^2 - 4R_{\mu\nu}R^{\mu\nu} + R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma}$ is the Gauss-Bonnet combination and $\omega(\phi)$ is the coupling function, while

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* Arpan's talk discussed the $\sigma = 0$ case.





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To allow much of this analysis to be confidently performed in the Jordan frame, as the conformal transformation of the Gauss-Bonnet term is rather complicated, we first rediscovered the known analytical predictions for standard Higgs inflation to leading order in slow-roll without appealing to a conformal transformation to the Einstein frame as is usually done.

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$$S^{J} = \int d^{4}x \sqrt{-g^{J}} \left[\frac{M_{p}^{2}}{2} \Omega^{2}(\phi) R^{J} - \frac{1}{2} g_{ab}^{J} \nabla^{a} \phi \nabla^{b} \phi - V(\phi) + \omega(\phi) R_{GB}^{2^{J}} \right]$$

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$$\Omega^{2}(\phi) = 1 + \frac{\sigma}{M_{p}^{2}}\phi^{2}, \quad V(\phi) = \frac{\lambda}{4}(\phi^{2} - \nu^{2})^{2}.$$

"We are interested in to see how this action change under the CT and what consequent dynamics would be apparent in the Einstein frame that otherwise does not come into sight in the Jordan frame"

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$$S^J = \int d^4x \sqrt{-g^J}$$

$$\frac{M_p^2}{2} \Omega^2(\phi) R^J -$$

We already know how this part transforms under the conformal transformation.

$$S = \int d^4 x \sqrt{-g} \left[\frac{M_P^2}{2} R - \frac{1}{2} g^{ab} \nabla_a s \nabla_b s - \frac{\lambda M_P^4}{4\sigma^2} \left(1 - e^{-\sqrt{\frac{2}{3}} \frac{s}{M_P}} \right)^2 \right]$$

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$$\frac{1}{2}g_{ab}^{J}\nabla^{a}\phi\nabla^{b}\phi - V(\phi) + \omega(\phi)R_{GB}^{2^{J}}$$



$$S^J = \int d^4x \sqrt{-g^J}$$

$$\begin{split} &\left[\frac{M_p^2}{2}\Omega^2(\phi)R^J - \frac{1}{2}g_{ab}^J\nabla^a\phi\,\nabla^b\phi - V(\phi) + \omega(\phi)R_{GB}^{2^J}\right] ,\\ &\text{We already know how this part transforms under the conformal transformation.} \\ &S = \int d^4x \sqrt{-g}\left[\frac{M_p^2}{2}R - \frac{1}{2}g^{ab}\nabla_a s\nabla_b s - \frac{\lambda M_p^4}{4\sigma^2}\left(1 - e^{-\sqrt{\frac{2}{3}}\frac{s}{M_p}}\right)^2\right] \end{split}$$

$$\frac{M_p^2}{2} \Omega^2(\phi) R^J - \frac{1}{2} g_{ab}^J \nabla^a \phi \nabla^b \phi - V(\phi) + \left[\omega(\phi) R_{GB}^{2^J} \right] ,$$

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We do not know how this term transforms under the conformal transformation.





$$S^J = \int d^4x \sqrt{-g^J}$$

$$\begin{split} & \left[\frac{M_p^2}{2}\Omega^2(\phi)R^J - \frac{1}{2}g_{ab}^J \nabla^a \phi \nabla^b \phi - V(\phi) + \omega(\phi)R_{GB}^{2^J}\right] , \\ & \text{We already know how this part transforms under the conformal transformation.} \\ & S = \int d^4x \sqrt{-g} \left[\frac{M_p^2}{2}R - \frac{1}{2}g^{ab}\nabla_a S\nabla_b S - \frac{\lambda M_p^4}{4\sigma^2} \left(1 - e^{-\sqrt{\frac{2}{3}}\frac{S}{M_p}}\right)^2\right] \end{split}$$

$$\frac{M_p^2}{2} \Omega^2(\phi) R^J - \frac{1}{2} g_{ab}^J \nabla^a \phi \nabla^b \phi - V(\phi) + \left[\omega(\phi) R_{GB}^{2^J} \right] ,$$

We already know how this part transforms under the conformal transformation.
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We do not know how this term transforms under the conformal transformation.

Let us focus on this term...





with non-minimal coupling + Gauss-Bonnet term in the Einstein frame

The Gauss-Bonnet part of the action reads:

$$S^J = \int d^4x$$

 $c_{\sqrt{-g^J}\omega(\phi)R_{GB}^{2^J}}$

Higgs Inflation with non-minimal coupling + Gauss-Bonnet term in the Einstein frame The Gauss-Bonnet part of the action reads: 0

$$S^{J} = \int d^{4}x \sqrt{-g^{J}}\omega(\phi) R_{GB}^{2^{J}}.$$

Under the $g_{ab}^{J} = \Omega^{-2}g_{ab}^{J}$, the Gauss-Bonnet term changes as $R_{GB}^{2^{J}} = \Omega^4 \left[R_{GB}^2 - 8\Omega^{-1} G_{ab} \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \right]$ $+8\Omega^{-2}\left(\nabla_{a}\nabla^{a}\Omega\nabla_{b}\nabla^{b}\Omega-\nabla_{b}\nabla_{a}\Omega\nabla^{b}\nabla^{a}\Omega\right)-$

*
$$R_{GB}^2 = R^2 - 4R_{\mu\nu}R^{\mu\nu} + R_{\mu\nu\rho\sigma}R^{\mu\nu\rho\sigma}$$

** $G_{ab} \equiv R_{ab} - g_{ab}R/2$

$$\begin{bmatrix} \nabla_a \Omega \nabla^a \Omega \\ -24\Omega^{-3} \nabla_a \Omega \nabla^a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \right)^2 \end{bmatrix}.$$

Grav. Cosmol. 10, 305 (2004)



Higgs Inflation with non-minimal coupling + Gauss-Bonnet term in the Einstein frame The Gauss-Bonnet part of the action reads: 0

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Under the $g_{ab}^{J} = \Omega^{-2}g_{ab}$, the Gauss-Bonnet term changes as $R_{GB}^{2'} = \Omega^4 \left[R_{GB}^2 - 8\Omega^{-1} G_{ab} \nabla^a \nabla^b \Omega - 4R \Omega^{-2} \right]$ $+8\Omega^{-2}\left(\nabla_{a}\nabla^{a}\Omega\nabla_{b}\nabla^{b}\Omega-\nabla_{b}\nabla_{a}\Omega\nabla^{b}\nabla^{a}\Omega\right)-$

Consequently, the action becomes

$$\begin{bmatrix} \nabla_a \Omega \nabla^a \Omega \\ -24\Omega^{-3} \nabla_a \Omega \nabla^a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \right)^2 \end{bmatrix}.$$
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 $\nabla_{a} \nabla^{a} \Omega \nabla_{b} \nabla^{b} \Omega - \nabla_{b} \nabla_{a} \Omega \nabla^{b} \nabla^{a} \Omega - 24 \Omega^{-3} \nabla_{a} \Omega \nabla^{a} \Omega \nabla_{b} \nabla^{b} \Omega + 24 \Omega^{-4} \left(\nabla_{a} \Omega \nabla^{a} \Omega \right)^{2} \right| .$





with non-minimal coupling + Gauss-Bonnet term in the Einstein frame

 $S = \left[d^4 x \sqrt{-g} \,\omega \left[R_{GB}^2 - 8\Omega^{-1} G_{ab} \nabla^a \nabla^b \Omega - 4R \Omega^{-2} \nabla_a \Omega \nabla^a \Omega + 8\Omega^{-2} \left(\nabla_a \nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \nabla_a \Omega \nabla^b \nabla^a \Omega \right) - 24\Omega^{-3} \nabla_a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \nabla^a \Omega \nabla^b \nabla^a \Omega \right)^2 \right].$

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 $S = \left[d^4 x \sqrt{-g} \,\omega \left[R_{GB}^2 - 8\Omega^{-1} G_{ab} \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \nabla_a \Omega \nabla^a \Omega + 8\Omega^{-2} \left(\nabla_a \nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \nabla_a \Omega \nabla^b \nabla^a \Omega \right) - 24\Omega^{-3} \nabla_a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \nabla^a \Omega \nabla^b \nabla^a \Omega \right)^2 \right].$

What happens if ω =const.?

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Do we expect to see any nontrivial effects of the "Gauss-Bonnet contributions" in the Einstein frame?

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with non-minimal coupling + Gauss-Bonnet term in the Einstein frame when $\omega = const$. $S = \omega \left[d^4 x \sqrt{-g} \left[R_{GB}^2 - 8\Omega^{-1} G_{ab} \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \nabla_a \Omega \nabla^a \Omega + 8\Omega^{-2} \left(\nabla_a \nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \nabla_a \Omega \nabla^b \nabla^a \Omega \right) - 24\Omega^{-3} \nabla_a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \nabla^a \Omega \nabla^b \nabla^a \Omega \right)^2 \right].$

What happens if $\omega = const.$?

 $-2\Omega^{-3}\nabla_a\Omega\nabla_b\Omega\nabla^a\nabla^b\Omega$

Do we expect to see any nontrivial effects of the "Gauss-Bonnet contributions" in the Einstein frame?

When $\omega = const.$, we obtain

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with non-minimal coupling + Gauss-Bonnet term in the Einstein frame when $\omega = const$. $S = \omega \left[d^4 x \sqrt{-g} \left[R_{GB}^2 - 8\Omega^{-1} G_{ab} \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \nabla_a \Omega \nabla^a \Omega + 8\Omega^{-2} \left(\nabla_a \nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \nabla_a \Omega \nabla^b \nabla^a \Omega \right) - 24\Omega^{-3} \nabla_a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \nabla^a \Omega \nabla^b \nabla^a \Omega \right)^2 \right].$

What happens if $\omega = const.$?

 $\mathbf{2}\nabla^a \mathbf{\Omega} - \frac{1}{2}\nabla^a (\nabla \mathbf{\Omega})^2$ $[\nabla_a, \nabla_b] V^c = - V^d \mathbf{R}^c{}_{dba}$



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What happens if $\omega = const.$?

$$S = \omega \int d^4x \sqrt{-g} \left[R_{GB}^2 - 8\Omega^{-1}G_{ab} \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \nabla_a \Omega \nabla^a \Omega + 8\Omega^{-2} \left(\nabla_a \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \nabla_a \Omega \nabla^a \Omega \right) \right]$$

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with non-minimal coupling + Gauss-Bonnet term in the Einstein frame when $\omega = const$. $7^{a}\Omega \nabla_{b}\nabla^{b}\Omega - \nabla_{b}\nabla_{a}\Omega \nabla^{b}\nabla^{a}\Omega - 24\Omega^{-3}\nabla_{a}\Omega \nabla^{b}\Omega + 24\Omega^{-4}\left(\nabla_{a}\Omega \nabla^{a}\Omega\right)^{2}.$

What happens if $\omega = const.$?

$\Omega \nabla_h \Omega \nabla^a \nabla^a \nabla^b$

$$S = \omega \int d^4x \sqrt{-g} \left[R_{GB}^2 - 8\Omega^{-1}G_{ab} \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \nabla_a \Omega \nabla^a \Omega + 8\Omega^{-2} \left(\nabla_a \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \nabla_a \Omega \nabla^a \Omega \right) \right]$$

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with non-minimal coupling + Gauss-Bonnet term in the Einstein frame when $\omega = const$. $\left| {}^{a}\Omega \nabla_{b}\nabla^{b}\Omega - \nabla_{b}\nabla_{a}\Omega \nabla^{b}\nabla^{a}\Omega \right) - 24\Omega^{-3}\nabla_{a}\Omega \nabla^{a}\Omega \nabla_{b}\nabla^{b}\Omega + 24\Omega^{-4} \left(\nabla_{a}\Omega \nabla^{a}\Omega \right)^{2} \right| .$ What happens if $\omega = const.?$ Do we expect to see any nontrivial effects of the "Gauss-Bonnet contributions" in the Einstein frame? $\Omega^{-2}G_{ab}\nabla^a\Omega\nabla^b\Omega$ $-8\omega \left[d^4x \sqrt{-g} \Omega^{-3} \nabla_a \Omega \nabla^a \Omega \nabla_b \nabla^b \Omega \right]$ $\nabla^a \mathbf{\Omega}$ $\nabla_a \Omega \nabla^a \Omega \nabla_b \nabla^b \Omega - 2\Omega^{-3} \nabla_a \Omega \nabla_b \Omega \nabla^a \nabla^b \Omega)$ $\int d^4x \sqrt{-g} \, \Omega^{-2} G^{ab} \, \nabla_a \Omega \, \nabla_b \Omega$



Do we expect to see any nontrivial effects of the "Gauss-Bonnet contributions" in the Einstein frame?

When $\omega = const.$, we obtain

$$S = \omega \int d^4x \sqrt{-g} \left[R_{GB}^2 - 16\Omega^{-3} \nabla_a \Omega \nabla_b \Omega \nabla^a \nabla^b \Omega - 8\Omega^{-3} \nabla_a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \right)^2 \right]$$

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$$\int S = \omega \int d^{4}x \sqrt{-g} \left[R_{GB}^{2} - 16\nabla_{a}\ln\Omega\nabla_{b}\ln\Omega\left(\Omega^{-1}\nabla^{a}\nabla^{b}\Omega\right) - 8\nabla_{a}\ln\Omega\nabla^{a}\ln\Omega\left(\Omega^{-1}\nabla_{b}\nabla^{b}\Omega\right) + 24\left(\nabla_{a}\ln\Omega\nabla^{a}\ln\Omega\nabla^{a}\ln\Omega\nabla^{a}\ln\Omega\nabla^{a}\ln\Omega\right) \right].$$



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Do we expect to see any nontrivial effects of the "Gauss-Bonnet contributions" in the Einstein frame?

When $\omega = const.$, we obtain

$$S = \omega \int d^{4}x \sqrt{-g} \left[R_{GB}^{2} - 16\Omega^{-3}\nabla_{a}\Omega\nabla_{b}\Omega\nabla^{a}\nabla^{b}\Omega - 8\Omega^{-3}\nabla_{a}\Omega\nabla_{b}\nabla^{b}\Omega + 24\Omega^{-4}\left(\nabla_{a}\Omega\nabla^{a}\Omega\right)^{2} \right].$$

$$\int d^{4}x \sqrt{-g} \left[R_{GB}^{2} - 16\nabla_{a}\ln\Omega\nabla_{b}\ln\Omega\left(\Omega^{-1}\nabla^{a}\nabla^{b}\Omega\right) - 8\nabla_{a}\ln\Omega\nabla^{a}\ln\Omega\left(\Omega^{-1}\nabla_{b}\nabla^{b}\Omega\right) + 24\left(\nabla_{a}\ln\Omega\nabla^{a}\ln\Omega\nabla^{a}\ln\Omega\nabla^{a}\ln\Omega\nabla^{a}\ln\Omega\right) \right].$$

$$Using \Omega^{-1}\nabla^{a}\nabla^{b}\Omega = \nabla^{a}\nabla^{b}\ln\Omega + \nabla^{a}\ln\Omega\nabla^{b}\ln\Omega$$

$$S = \omega \int d^4x \sqrt{-g} \left[R_{GB}^2 - 16 \nabla_a \ln \Omega \nabla_b \ln \Omega \nabla^a \nabla \nabla_b \right]$$

 $\nabla^{p}\ln\Omega - 8\nabla_{a}\ln\Omega\nabla^{a}\ln\Omega\nabla_{b}\nabla^{p}\ln\Omega$



What happens if ω =const.?

Do we expect to see any nontrivial effects of the "Gauss-Bonnet contributions" in the Einstein frame?

When ω =const., we obtain

$$S = \omega \int d^{4}x \sqrt{-g} \left[R_{GB}^{2} - 16\Omega^{-3} \nabla_{a}\Omega \nabla_{b}\Omega \nabla^{a}\nabla^{b}\Omega - 8\Omega^{-3} \nabla_{a}\Omega \nabla_{b}\nabla^{b}\Omega + 24\Omega^{-4} \left(\nabla_{a}\Omega \nabla^{a}\Omega\right)^{2} \right].$$

$$\int d^{4}x \sqrt{-g} \left[R_{GB}^{2} - 16 \nabla_{a}\ln\Omega \nabla_{b}\ln\Omega \left(\Omega^{-1}\nabla^{a}\nabla^{b}\Omega\right) - 8 \nabla_{a}\ln\Omega \nabla^{a}\ln\Omega \left(\Omega^{-1}\nabla_{b}\nabla^{b}\Omega\right) + 24 \left(\nabla_{a}\ln\Omega \nabla^{a}\ln\Omega \nabla^{a}\ln\Omega$$



Do we expect to see any nontrivial effects of the "Gauss-Bonnet contributions" in the Einstein frame?

When $\omega = const.$, we obtain



The Gauss-Bonnet coupling must be a function of a scalar field, i.e., $\omega(\phi)!$



 $S = \left[d^4 x \sqrt{-g} \,\omega(\phi) \left[R_{GB}^2 - 8\Omega^{-1} G_{ab} \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \nabla_a \Omega \nabla^a \Omega + 8\Omega^{-2} \left(\nabla_a \nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \nabla_a \Omega \nabla^b \nabla^a \Omega \right) - 24\Omega^{-3} \nabla_a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \nabla^a \Omega \nabla^b \nabla^a \Omega \right)^2 \right].$

This time $\omega(\phi)$ is a function of a scalar field ϕ .



*
$$-8\int d^4x \sqrt{-g} \omega \left[\Omega^{-1}G_{ab} \nabla^a \nabla^b \Omega + \frac{1}{2}g_{ab}R\Omega^{-2} \nabla^a \Omega \nabla^b \Omega \right]$$

**
$$G_{ab} \equiv R_{ab} - g_{ab}R/2$$

 $= -8 \left[d^4 x \sqrt{-g} \left[\omega \Omega^{-2} R_{ab} \nabla^a \Omega \nabla^b \Omega - \Omega^{-1} G_{ab} \nabla^a \omega \nabla^b \Omega \right] \right]$



$$\begin{split} & \textbf{Higgs Inflation} \\ with non-minimal coupling + Gauss-Bonnet term in the Einstein fraction for the formula form$$

$\left(abla_a \Omega abla^a \Omega ight)^2 ight].$



$$\begin{split} & \textbf{Higgs Inflation} \\ with non-minimal coupling + Gauss-Bonnet term in the Einstein framework \\ & s = \int d^{4}x \sqrt{-g} \, \omega(\phi) \Big[R_{GB}^{2} - 8\Omega^{-1}G_{ab} \nabla^{a}\nabla^{b}\Omega - 4R\Omega^{-2} \nabla_{a}\Omega \nabla^{a}\Omega + 8\Omega^{-2} (\nabla_{a}\nabla^{a}\Omega \nabla_{b}\nabla^{b}\Omega - \nabla_{b}\nabla_{a}\Omega \nabla^{b}\nabla^{a}\Omega) - 24\Omega^{-3}\nabla_{a}\Omega \nabla^{a}\Omega \nabla_{b}\nabla^{b}\Omega + 24\Omega^{-4} (\nabla_{a}\nabla^{a}\Omega \nabla_{b}\nabla^{b}\Omega - \nabla_{b}\nabla_{a}\Omega \nabla^{b}\Omega - 24\Omega^{-3}\nabla_{a}\Omega \nabla^{a}\Omega \nabla_{b}\nabla^{b}\Omega + 24\Omega^{-4} (\nabla_{a}\nabla^{a}\nabla^{a}\nabla^{b}\nabla^{a}\Omega - \nabla_{b}\nabla_{a}\Omega \nabla^{b}\Omega - \nabla_{b}\nabla_{a}\Omega \nabla^{b}\Omega - \nabla_{b}\nabla_{a}\Omega \nabla^{b}\Omega - 24\Omega^{-3}\nabla_{a}\Omega \nabla^{b}\Omega - \Omega^{-1}G_{ab}\nabla^{a}\Omega \nabla^{b}\Omega + 24\Omega^{-4} (\nabla_{a}\nabla^{a}\nabla^{a}\nabla^{b}\nabla^{a}\Omega - \nabla_{b}\nabla_{a}\Omega \nabla^{b}\Omega - \nabla_{b}\nabla_{a}\Omega \nabla^{b}\Omega - \nabla_{b}\nabla_{a}\Omega \nabla^{b}\Omega - \Omega^{-1}G_{ab}\nabla^{a}\omega \nabla^{b}\Omega - \nabla_{b}\nabla_{a}\Omega \nabla^{b}\nabla^{a}\Omega - \Omega^{-1}\nabla_{a}\Omega (\nabla^{a}\Omega \nabla_{b}\nabla^{b}\Omega - \nabla_{b}\Omega \nabla^{a}\nabla^{b}\Omega) \\ &= 8\int d^{4}x \sqrt{-g} \left[\omega\Omega^{-2}R_{ab}\nabla^{a}\Omega \nabla^{b}\Omega - \omega\Omega^{-2} (\omega^{-1}\nabla_{a}\omega - 2\Omega^{-1}\nabla_{a}\Omega) (\nabla^{a}\Omega \nabla_{b}\nabla^{b}\Omega - \nabla_{b}\Omega \nabla^{a}\nabla^{b}\Omega) \right] \end{split}$$

$\left(abla_a \Omega abla^a \Omega ight)^2 ight].$



Higgs Inflation $S = \left[d^4 x \sqrt{-g} \,\omega(\phi) \left[R_{GB}^2 - 8\Omega^{-1} G_{ab} \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \nabla_a \Omega \nabla^a \Omega + 8\Omega^{-2} \left(\nabla_a \nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \nabla_a \Omega \nabla^b \nabla^a \Omega \right) - 24\Omega^{-3} \nabla_a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \nabla^a \Omega \nabla^b \nabla^a \Omega \right)^2 \right].$ * $-8\int d^4x \sqrt{-g}\,\omega \left[\Omega^{-1}G_{ab}\nabla^a\nabla^b\Omega + \frac{1}{2}g_{ab}R\Omega^{-2}\nabla^a\Omega\nabla^b\Omega\right] = -\int d^4x \sqrt{-g}\left[\omega\Omega^{-2}R_{ab}\nabla^a\Omega\nabla^b\Omega - \Omega^{-1}G_{ab}\nabla^a\omega\nabla^b\Omega\right]$ * $8\int d^4x \sqrt{-g}\omega \Omega^{-2} \left(\nabla_a \nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \nabla_a \Omega \nabla^b \nabla^a \Omega \right)$ $S = \left[d^{4}x \sqrt{-g} \left[\omega R_{GB}^{2} + 8\Omega^{-1}G_{ab} \nabla^{a}\omega \nabla^{b}\Omega - 8\omega\Omega^{-2} \left(\omega^{-1} \nabla_{a}\omega - 2\Omega^{-1} \nabla_{a}\Omega \right) \left(\nabla^{a}\Omega \nabla_{b} \nabla^{b}\Omega - \nabla_{b}\Omega \nabla^{a}\nabla^{b}\Omega \right) - 24\omega\Omega^{-3} \nabla_{a}\Omega \nabla_{b} \nabla^{b}\Omega + 24\omega\Omega^{-4} \left(\nabla_{a}\Omega \nabla^{a}\Omega \nabla^{a}\Omega \right)^{2} \right]$

with non-minimal coupling + Gauss-Bonnet term in the Einstein frame







Higgs Inflation $S = \left[d^4 x \sqrt{-g} \,\omega(\phi) \left[R_{GB}^2 - 8\Omega^{-1} G_{ab} \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \nabla_a \Omega \nabla^a \Omega + 8\Omega^{-2} \left(\nabla_a \nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \nabla_a \Omega \nabla^b \nabla^a \Omega \right) - 24\Omega^{-3} \nabla_a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \nabla^a \Omega \nabla^b \nabla^a \Omega \right)^2 \right] \,.$ * $-8\int d^4x \sqrt{-g}\,\omega \left[\Omega^{-1}G_{ab}\nabla^a\nabla^b\Omega + \frac{1}{2}g_{ab}R\Omega^{-2}\nabla^a\Omega\nabla^b\Omega\right] = \int d^4x \sqrt{-g}\left[\omega\Omega^{-2}R_{ab}\nabla^a\Omega\nabla^b\Omega - \Omega^{-1}G_{ab}\nabla^a\omega\nabla^b\Omega\right]$ * $8\int d^4x \sqrt{-g}\omega \Omega^{-2} \left(\nabla_a \nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \nabla_a \Omega \nabla^b \nabla^a \Omega \right)$ $S = \left[d^{4}x \sqrt{-g} \left[\omega R_{GB}^{2} + 8\Omega^{-1}G_{ab} \nabla^{a}\omega \nabla^{b}\Omega - 8\omega\Omega^{-2} \left(\omega^{-1} \nabla_{a}\omega - 2\Omega^{-1} \nabla_{a}\Omega \right) \left(\nabla^{a}\Omega \nabla_{b} \nabla^{b}\Omega - \nabla_{b}\Omega \nabla^{a}\nabla^{b}\Omega \right) - 24\omega\Omega^{-3} \nabla_{a}\Omega \nabla_{b}\nabla^{b}\Omega + 24\omega\Omega^{-4} \left(\nabla_{a}\Omega \nabla^{a}\Omega \nabla^{a}\Omega \right)^{2} \right]$ If $\omega = \alpha \Omega^2$, $\left(\omega^{-1} \nabla_a \omega - 2 \Omega^{-1} \nabla_a \Omega \right) = 0$.

with non-minimal coupling + Gauss-Bonnet term in the Einstein frame







Higgs Inflation $S = \left[d^4 x \sqrt{-g} \,\omega(\phi) \left[R_{GB}^2 - 8\Omega^{-1} G_{ab} \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \nabla_a \Omega \nabla^a \Omega + 8\Omega^{-2} \left(\nabla_a \nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \nabla_a \Omega \nabla^b \nabla^a \Omega \right) - 24\Omega^{-3} \nabla_a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \nabla^a \Omega \nabla^b \nabla^a \Omega \right)^2 \right] \,.$ * $-8\int d^4x \sqrt{-g}\,\omega \left[\Omega^{-1}G_{ab}\nabla^a\nabla^b\Omega + \frac{1}{2}g_{ab}R\Omega^{-2}\nabla^a\Omega\nabla^b\Omega\right] = \int d^4x \sqrt{-g}\left[\omega\Omega^{-2}R_{ab}\nabla^a\Omega\nabla^b\Omega - \Omega^{-1}G_{ab}\nabla^a\omega\nabla^b\Omega\right]$ * $8\int d^4x \sqrt{-g}\omega \Omega^{-2} \left(\nabla_a \nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \nabla_a \Omega \nabla^b \nabla^a \Omega \right)$ $= 8 \left[d^4 x \sqrt{-g} \left[\omega \Omega^{-2} R_{ab} \nabla^a \Omega \nabla^b \Omega - \omega \Omega^{-2} \left(\omega^{-1} \sqrt{-2\Omega^{-1} \nabla_a \Omega} \right) \left(\nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \Omega \nabla^a \nabla^b \Omega \right) \right]$ $S = \left[d^{4}x \sqrt{-g} \left[\omega R_{GB}^{2} + 8\Omega^{-1}G_{ab} \nabla^{a}\omega \nabla^{b}\Omega - 8\omega\Omega^{-2} \left(\omega^{-1} \nabla_{a}\omega - 2\Omega^{-1} \nabla_{a}\Omega \right) \left(\nabla^{a}\Omega \nabla_{b} \nabla^{b}\Omega - \nabla_{b}\Omega \nabla^{a}\nabla^{b}\Omega \right) - 24\omega\Omega^{-3} \nabla_{a}\Omega \nabla_{b}\nabla^{b}\Omega + 24\omega\Omega^{-4} \left(\nabla_{a}\Omega \nabla^{a}\Omega \nabla^{a}\Omega \right)^{2} \right]$ If $\omega = \alpha \Omega^2$, $\left(\omega^{-1} \nabla_a \omega - 2 \Omega^{-1} \nabla_a \Omega \right) = 0$. $S = \left[d^4 x \sqrt{-g} \alpha \Omega^2 \left[R_{GB}^2 + 16 \Omega^{-2} G_{ab} \nabla^a \Omega \nabla^b \Omega - 24 \Omega^{-3} \nabla_a \Omega \nabla^a \Omega \nabla_b \nabla^b \Omega + 24 \Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \right)^2 \right]$



with non-minimal coupling + Gauss-Bonnet term in the Einstein frame





 $S = \left[d^4 x \sqrt{-g} \,\omega(\phi) \left[R_{GB}^2 - 8\Omega^{-1} G_{ab} \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \nabla_a \Omega \nabla^a \Omega + 8\Omega^{-2} \left(\nabla_a \nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \nabla_a \Omega \nabla^b \nabla^a \Omega \right) - 24\Omega^{-3} \nabla_a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \nabla^b \Omega \right)^2 \right].$

$$S = \int d^4x \sqrt{-g} \alpha \Omega^2 \left[R_{GB}^2 + 16G_{ab} \nabla^a \ln \Omega \nabla^b \ln \Omega - \frac{1}{2} \nabla^a \ln \Omega \nabla^b \ln \Omega \right]$$

$$S = \int d^4x \sqrt{-g} \alpha \Omega^2 \left[R_{GB}^2 + 16\Omega^{-2} G_{ab} \nabla^a \Omega \nabla^b \right]$$

 $^{b}\Omega - 24\Omega^{-3}\nabla_{a}\Omega\nabla^{a}\Omega\nabla_{b}\nabla^{b}\Omega + 24\Omega^{-4}\left(\nabla_{a}\Omega\nabla^{a}\Omega\right)^{2}$

 $-24\nabla_{a}\ln\Omega\nabla^{a}\ln\Omega\left(\Omega^{-1}\nabla_{b}\nabla^{b}\Omega\right)+24\left(\nabla_{a}\ln\Omega\nabla^{a}\ln\Omega\right)^{2}$

















 $S = \left[d^4 x \sqrt{-g} \,\omega(\phi) \left[R_{GB}^2 - 8\Omega^{-1} G_{ab} \nabla^a \nabla^b \Omega - 4R\Omega^{-2} \nabla_a \Omega \nabla^a \Omega + 8\Omega^{-2} \left(\nabla_a \nabla^a \Omega \nabla_b \nabla^b \Omega - \nabla_b \nabla_a \Omega \nabla^b \nabla^a \Omega \right) - 24\Omega^{-3} \nabla_a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \nabla^a \Omega \nabla^b \nabla^a \Omega \right)^2 \right].$

$$S = \int d^4x \sqrt{-g} \alpha \Omega^2 \left[R_{GB}^2 + 16\Omega^{-2} G_{ab} \nabla^a \Omega \nabla^b \Omega - 24\Omega^{-3} \nabla_a \Omega \nabla^a \Omega \nabla_b \nabla^b \Omega + 24\Omega^{-4} \left(\nabla_a \Omega \nabla^a \Omega \right)^2 \right]$$

















$$S = \int d^{4}x \sqrt{-g} \, \omega(\phi) \left[R_{GB}^{2} - 8\Omega^{-1}G_{ab} \nabla^{a}\nabla^{b}\Omega - 4R\Omega^{-2}\nabla_{a}\Omega\nabla^{a}\Omega + 8\Omega^{-2} (\nabla_{a}\nabla^{a}\Omega\nabla_{b}\nabla^{b}\Omega - \nabla_{b}\nabla_{a}\Omega\nabla^{b}\nabla^{a}\Omega) - 24\Omega^{-3}\nabla_{a}\Omega\nabla^{a}\Omega\nabla_{b}\nabla^{b}\Omega + 24\Omega^{-4} (\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\nabla^{a}\Omega) - 24\Omega^{-3}\nabla_{a}\Omega\nabla^{a}\Omega\nabla_{b}\nabla^{b}\Omega + 24\Omega^{-4} (\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega) - 24\Omega^{-3}\nabla_{a}\Omega\nabla^{a}\Omega\nabla_{b}\nabla^{b}\Omega + 24\Omega^{-4} (\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega) - 24\Omega^{-3}\nabla_{a}\Omega\nabla^{a}\Omega\nabla_{b}\nabla^{b}\Omega + 24\Omega^{-4} (\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega) - 24\Omega^{-3}\nabla_{a}\Omega\nabla^{a}\Omega\nabla_{b}\nabla^{b}\Omega + 24\Omega^{-4} (\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega) + 24\Omega^{-4}(\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega) + 24\Omega^{-4}(\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega) + 24\Omega^{-4}(\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega) + 24\Omega^{-4}(\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega) + 24\Omega^{-4}(\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega) + 24\Omega^{-4}(\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega) + 24\Omega^{-4}(\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega\nabla^{b}\Omega) + 24\Omega^{-4}(\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega\nabla^{b}\Omega) + 24\Omega^{-4}(\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega\nabla^{b}\Omega\nabla^{b}\Omega\nabla^{b}\Omega\nabla^{b}\Omega\nabla^{b}\Omega\nabla^{b}\Omega) + 24\Omega^{-4}(\nabla_{a}\Omega\nabla^{a}\Omega\nabla^{a}\Omega\nabla^{b}\Omega\nabla^$$

$$\omega = \alpha \Omega^{2} \qquad S = \int d^{4}x \sqrt{-g} \times \alpha e^{\sqrt{\frac{2}{3}} \frac{x}{M_{P}}} \left[R_{GB}^{2} + \frac{8}{3M_{P}^{2}} G_{ab} \nabla^{a} s \nabla^{b} s - \frac{1}{M_{P}^{3}} \sqrt{\frac{8}{3}} \nabla^{b} \nabla_{b} s \nabla_{a} s \nabla^{a} s \right]$$

$$\omega = \alpha \Omega^{2} \qquad S = \int d^{4}x \sqrt{-g} \alpha \Omega^{2} \left[R_{GB}^{2} + 16G_{ab} \nabla^{a} \ln \Omega \nabla^{b} \ln \Omega - 24 \nabla_{a} \ln \Omega \nabla^{a} \ln \Omega \nabla_{b} \nabla^{b} \ln \Omega \right]$$

$$\omega = \alpha \Omega^{2} \qquad S = \int d^{4}x \sqrt{-g} \alpha \Omega^{2} \left[R_{GB}^{2} + 16G_{ab} \nabla^{a} \ln \Omega \nabla^{b} \ln \Omega - 24 \nabla_{a} \ln \Omega \nabla^{a} \ln \Omega \nabla_{b} \nabla^{b} \ln \Omega \right]$$

$$\omega = \alpha \Omega^{2} \qquad U \sin \alpha \Omega^{-1} \nabla^{a} \nabla^{b} \Omega = \nabla^{a} \nabla^{b} \ln \Omega + \nabla^{a} \ln \Omega \nabla^{b} \ln \Omega$$

$$S = \int d^{4}x \sqrt{-g} \alpha \Omega^{2} \left[R_{GB}^{2} + 16G_{ab} \nabla^{a} \ln \Omega \nabla^{b} \ln \Omega - 24 \nabla_{a} \ln \Omega \nabla^{a} \ln \Omega (\Omega^{-1} \nabla_{b} \nabla^{b} \Omega) + 24 (\nabla_{a} \ln \Omega \nabla^{a} \ln \Omega)^{2} \right]$$

$$S = \int d^{4}x \sqrt{-g} \alpha \Omega^{2} \left[R_{GB}^{2} + 16\Omega^{-2}G_{ab} \nabla^{a} \Omega \nabla^{b} \Omega - 24\Omega^{-3} \nabla_{a} \Omega \nabla^{a} \Omega \nabla_{b} \nabla^{b} \Omega + 24\Omega^{-4} (\nabla_{a} \Omega \nabla^{a} \Omega)^{2} \right]$$





Combining with the minimal coupling case,

$$S = \int d^4x \sqrt{-g} \left[\frac{M_P^2}{2} R - \frac{1}{2} (\partial s)^2 - V(s) - \frac{1}{2} \xi(s) \left(c_1 R_{GB}^2 + \frac{c_2}{M_P^2} G_{ab} \nabla^a s \nabla^b s + \frac{c_3}{M_P^3} \nabla_a s \nabla^a s \nabla^b \nabla_b s \right) \right],$$

where
$$V(s) \equiv \frac{\lambda M_p^4}{4\sigma^2} \left(1 - e^{-\sqrt{\frac{2}{3}}\frac{s}{M_p}}\right)^2$$
, $\xi(s) \equiv -2\alpha e^{\sqrt{\frac{2}{3}}\frac{s}{M_p}}$, $c_1 = 1$, $c_2 = 8/3$, $c_3 = -\sqrt{8/3}$.
 $\lambda/\sigma^2 \simeq 5 \times 10^{-10}$

arXiv:2308.00897



Combining with the minimal coupling case,

$$S = \int d^4x \sqrt{-g} \left[\frac{M_P^2}{2} R - \frac{1}{2} (\partial s)^2 - V(s) - \frac{1}{2} \xi(s) \left(c_1 R_{GB}^2 + \frac{c_2}{M_P^2} G_{ab} \nabla^a s \nabla^b s + \frac{c_3}{M_P^3} \nabla_a s \nabla^a s \nabla^b \nabla_b s \right) \right],$$

where $V(s) \equiv \frac{\lambda M_P^4}{4\sigma^2} \left(1 - e^{-\sqrt{\frac{2}{3}} \frac{s}{M_P}} \right)^2$, $\xi(s) \equiv -2\alpha e^{\sqrt{\frac{2}{3}} \frac{s}{M_P}}$, $c_1 = 1$, $c_2 = 8/3$, $c_3 = -\sqrt{8/3}$.

We consider an action in Eq. (45) with the following additional corrections in the action [20,21]:

$$L_{(c)} = -\frac{1}{2}\xi(\phi)[c_1R_{GB}^2 + c_2G^{ab}\phi_{,a}\phi_{,b} + c_3\Box\phi\phi^{,c}\phi_{,c} + c_4(\phi^{,c}\phi_{,c})^2],$$

arXiv:2308.00897



F. String corrections

JAI-CHAN HWANG AND HYERIM NOH PHYSICAL REVIEW D 71, 063536 (2005)





Combining with the minimal coupling case,

$$S = \int d^{4}x \sqrt{-g} \left[\frac{M_{p}^{2}}{2}R - \frac{1}{2}(\partial s)^{2} - V(s) - \frac{1}{2}\xi(s) \left(c_{1}R_{GB}^{2} + \frac{c_{2}}{M_{p}^{2}}G_{ab}\nabla^{a}s\nabla^{b}s + \frac{c_{3}}{M_{p}^{3}}\nabla_{a}s\nabla^{a}s\nabla^{b}\nabla_{b}s \right) \right],$$
where $V(s) \equiv \frac{\lambda M_{p}^{4}}{4\sigma^{2}} \left(1 - e^{-\sqrt{\frac{2}{3}}\frac{s}{M_{p}}} \right)^{2}, \quad \xi(s) \equiv -2\alpha e^{\sqrt{\frac{2}{3}}\frac{s}{M_{p}}}, \quad c_{1} = 1, \quad c_{2} = 8/3, \quad c_{3} = -\sqrt{8/3}.$

 $V(s)$

 $V(s)$



arXiv:2308.00897





Combining with the minimal coupling case,

$$S = \int d^4x \sqrt{-g} \left[\frac{M_P^2}{2} R - \frac{1}{2} (\partial s)^2 - V(s) - \frac{1}{2} \xi(s) \left(c_1 R_{GB}^2 + \frac{c_2}{M_P^2} G_{ab} \nabla^a s \nabla^b s + \frac{c_3}{M_P^3} \nabla_a s \nabla^a s \nabla^b \nabla_b s \right) \right],$$

where $V(s) \equiv \frac{\lambda M_P^4}{2} \left(1 - e^{-\sqrt{\frac{2}{3}} \frac{s}{M_P}} \right)^2$, $\xi(s) \equiv -2\alpha e^{\sqrt{\frac{2}{3}} \frac{s}{M_P}}$, $c_1 = 1$, $c_2 = 8/3$, $c_3 = -\sqrt{8/3}$.

In the flat FRW universe with metric: $ds^2 = -dt^2 + a(t)^2 \delta_{ij} dx^i dx^j$, the background equations of motion read:

$$\begin{split} 3M_p^2 H^2 &= \frac{1}{2}\dot{s}^2 + V + 12c_1\dot{\xi}H^3 - \frac{9}{2}\frac{c_2}{M_p^2}\xi\dot{s}^2H^2 + \frac{1}{2}\frac{c_3}{M_p^3}(\dot{\xi} - 6\xiH)\dot{s}^3, \\ M_p^2(2\dot{H} + 3H^2) &= -\frac{1}{2}\dot{s}^2 + V + 4c_1\left[\ddot{\xi}H^2 + 2\dot{\xi}H(\dot{H} + H^2)\right] - \frac{1}{2}\frac{c_2}{M_p^2}\dot{s}\left[\xi\dot{s}(2\dot{H} + 3H^2) + 4\xi\ddot{s}H + 2\dot{\xi}\dot{s}H\right] - \frac{1}{2}\frac{c_3}{M_p^3}\dot{s}^2(2\xi\ddot{s} + \dot{\xi}\dot{s}), \\ \ddot{s} + 3H\dot{s} + V_{,s} &= -12c_1\xi_{,s}H^2(\dot{H} + H^2) + \frac{3}{2}\frac{c_2}{M_p^2}\left[H^2(\dot{\xi}\dot{s} + 2\xi\ddot{s}) + 2H\xi\dot{s}(2\dot{H} + 3H^2)\right] - \frac{1}{2}\frac{c_3}{M_p^3}\dot{s}\left[\ddot{\xi}\dot{s} + 3\dot{\xi}\ddot{s} - 6\xi(\dot{H}\dot{s} + 2H\ddot{s} + 3H^2\dot{s})\right] \end{split}$$

$$\begin{split} 3M_p^2 H^2 &= \frac{1}{2}\dot{s}^2 + V + 12c_1\dot{\xi}H^3 - \frac{9}{2}\frac{c_2}{M_p^2}\xi\dot{s}^2H^2 + \frac{1}{2}\frac{c_3}{M_p^3}(\dot{\xi} - 6\xiH)\dot{s}^3, \\ M_p^2(2\dot{H} + 3H^2) &= -\frac{1}{2}\dot{s}^2 + V + 4c_1\left[\ddot{\xi}H^2 + 2\dot{\xi}H(\dot{H} + H^2)\right] - \frac{1}{2}\frac{c_2}{M_p^2}\dot{s}\left[\xi\dot{s}(2\dot{H} + 3H^2) + 4\xi\ddot{s}H + 2\dot{\xi}\dot{s}H\right] - \frac{1}{2}\frac{c_3}{M_p^3}\dot{s}^2(2\xi\ddot{s} + \dot{\xi}\dot{s}), \\ \ddot{s} + 3H\dot{s} + V_{,s} &= -12c_1\xi_{,s}H^2(\dot{H} + H^2) + \frac{3}{2}\frac{c_2}{M_p^2}\left[H^2(\dot{\xi}\dot{s} + 2\xi\ddot{s}) + 2H\xi\dot{s}(2\dot{H} + 3H^2)\right] - \frac{1}{2}\frac{c_3}{M_p^3}\dot{s}\left[\ddot{\xi}\dot{s} + 3\dot{\xi}\ddot{s} - 6\xi(\dot{H}\dot{s} + 2H\ddot{s} + 3H^2\dot{s})\right] \end{split}$$

P

arXiv:2308.00897



The solutions to the background equations of motion:

$$V(s) \equiv \frac{\lambda M_p^4}{4\sigma^2} \left(1 - e^{-\sqrt{\frac{2}{3}}\frac{s}{M_p}}\right)^2, \quad \xi(s) \equiv -2\alpha e^{\sqrt{\frac{2}{3}}\frac{s}{M_p}}$$

$$3M_p^2 H^2 = \frac{1}{2}\dot{s}^2 + V + 12c_1\dot{\xi}H^3 - \frac{9}{2}\frac{c_2}{M_p^2}\xi\dot{s}^2H^2 + \frac{1}{2}\frac{c_3}{M_p^3}(\dot{\xi} - 6\xi H)$$
$$M^2(2\dot{H} + 3H^2) - \frac{1}{2}\dot{s}^2 + V + 4c_1\left[\ddot{\xi}H^2 + 2\dot{\xi}H(\dot{H} + H^2)\right] - \frac{1}{2}$$



$$\frac{c_2}{M_p^2} \dot{s} \left[\xi \dot{s} (2\dot{H} + 3H^2) + 4\xi \ddot{s} H + 2\dot{\xi} \dot{s} H \right] - \frac{1}{2} \frac{c_3}{M_p^3} \dot{s}^2 (2\xi \ddot{s} + \dot{\xi} \dot{s}),$$

 $\ddot{s} + 3H\dot{s} + V_{,s} = -12c_1\xi_{,s}H^2(\dot{H} + H^2) + \frac{3}{2}\frac{c_2}{M_p^2} \left[H^2(\dot{\xi}\dot{s} + 2\xi\ddot{s}) + 2H\xi\dot{s}(2\dot{H} + 3H^2) \right] - \frac{1}{2}\frac{c_3}{M_p^3}\dot{s} \left[\ddot{\xi}\dot{s} + 3\dot{\xi}\ddot{s} - 6\xi(\dot{H}\dot{s} + 2H\ddot{s} + 3H^2\dot{s}) \right].$

are simplified as

where
$$\delta(s) \equiv \frac{\mathscr{B} - \sqrt{\mathscr{B}^2 - 4\mathscr{A}\mathscr{C}}}{2\mathscr{A}V_{,s}} - 1$$
,

with
$$\mathscr{A} \equiv \frac{c_3}{M_p^3} \xi$$
, $\mathscr{B} \equiv 1 - \frac{3c_2}{M_p^2} \xi H^2$, $\mathscr{C} \equiv V_{,s} + 12c_1 \xi_{,s} H^4$.

The number of e-folds: $N \equiv \int_{t_i}^{t_e} H dt = \int_{s_i}^{s_e} \frac{H}{\dot{s}} ds$, where $\frac{\dot{s}}{H} \simeq -M_p^2 \frac{V_{,s}}{V}(1+\delta)$,

with non-minimal coupling + Gauss-Bonnet term in the Einstein frame In the context of slow-roll inflation, *i.e.*, $V \gg \dot{s}^2$, $\ddot{s} \ll 3H\dot{s}$, and $\dot{\xi}/(2\xi H) \ll 1$, the equations of motion





Combining with the minimal coupling case,

$$S = \int d^{4}x \sqrt{-g} \left[\frac{M_{P}^{2}}{2}R - \frac{1}{2}(\partial s)^{2} - V(s) - \frac{1}{2}\xi(s) \left(c_{1}R_{GB}^{2} + \frac{c_{2}}{M_{P}^{2}}G_{ab}\nabla^{a}s\nabla^{b}s + \frac{c_{3}}{M_{P}^{3}}\nabla_{a}s\nabla^{a}s\nabla^{b}\nabla_{b}s \right) \right],$$

where $V(s) \equiv \frac{\lambda M_{P}^{4}}{4\sigma^{2}} \left(1 - e^{-\sqrt{\frac{2}{3}}\frac{s}{M_{P}}} \right)^{2}, \quad \xi(s) \equiv -2\alpha e^{\sqrt{\frac{2}{3}}\frac{s}{M_{P}}}, \quad c_{1} = 1, \quad c_{2} = 8/3, \quad c_{3} = -\sqrt{8/3}.$

$$L_{(c)} = -\frac{1}{2}\xi(\phi)[c_1R_{GB}^2 + c_2G^{ab}\phi_{,a}\phi_{,b} + c_3\Box\phi\phi^{,c}\phi_{,c} + c_4(\phi^{,c}\phi_{,c})^2],$$

F. String corrections

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We consider an action in Eq. (45) with the following additional corrections in the action [20,21]:



 $n_S - 1 = 2(2\epsilon_1 - \epsilon_2 - \epsilon_3), \quad n_T = 2(\epsilon_1 - \epsilon_5), \quad r = 1$

$$\epsilon_1 \equiv \frac{\dot{H}}{H^2}, \ \epsilon_2 \equiv \frac{\ddot{s}}{H\dot{s}}, \ \epsilon_3 \equiv \frac{\dot{E}}{2EH}, \ \epsilon_4 \equiv \frac{Q_a}{2H(2M_p^2 + Q_b)}, \ \epsilon_5 \equiv \frac{\dot{Q}_t}{2Q_t H}, \ \text{where} \ E \equiv \frac{1}{\dot{s}^2} \left(\dot{s}^2 + \frac{3Q_a^2}{2M_p^2 + Q_b} + Q_c \right),$$

$$Q_a \equiv -4c_1 \dot{\xi} H^2 + \frac{2c_2}{M_p^2} \xi \dot{s}^2 H + \frac{c_3}{M_p^3} \xi \dot{s}^3, Q_b \equiv -8c_1 \dot{\xi} H + \frac{c_2}{M_p^2} \xi \dot{s}^2, Q_c \equiv -\frac{3c_2}{M_p^2} \xi \dot{s}^2 H^2 + \frac{2c_3}{M_p^3} \dot{s}^3 (\dot{\xi} - 3\xi H),$$

$$Q_{d} \equiv -\frac{2c_{2}}{M_{p}^{2}}\xi\dot{s}^{2}\dot{H} - \frac{2c_{3}}{M_{p}^{3}}\dot{s}^{2}(\dot{\xi}\dot{s} + \xi\ddot{s} - \xi\dot{s}H), \ Q_{e} \equiv -16c_{1}\dot{\xi}\dot{H} + \frac{2c_{2}}{M_{p}^{2}}\dot{s}(\dot{\xi}\dot{s} + 2\xi\ddot{s} - 2\xi\dot{s}H) - \frac{4c_{3}}{M_{p}^{3}}\xi\dot{s}^{3}, \ Q_{f} \equiv 8c_{1}(\ddot{\xi} - \dot{\xi}H) + \frac{2c_{2}}{M_{p}^{2}}\xi\dot{s}^{2}, \\ Q_{t} \equiv 1 + \frac{2c_{2}}{M_{p}^{2}}\xi\dot{s}^{2}, \ Q_{t} \equiv 1 + \frac{2c_{2}}{M_{p}^{2}}\xi\dot{s}^{2} + \frac{2c_{3}}{M_{p}^{2}}\xi\dot{s}^{2} + \frac{2c_{3$$

$$c_A^2 \equiv 1 + \frac{Q_d + \frac{Q_a}{2M_p^2 + Q_b}Q_e + \left(\frac{Q_a}{2M_p^2 + Q_b}\right)^2 Q_f}{\dot{s}^2 + \frac{3Q_a^2}{2M_p^2 + Q_b} + Q_c}, \qquad c_T^2 \equiv 1 - \frac{Q_f}{2M_p^2 + Q_b},$$

$$6 \left| \frac{1}{Q_t} \left(\frac{c_A}{c_T} \right)^3 \left(\epsilon_1 - \frac{1}{4M_p^2 H^2} \left(2Q_c + Q_d - HQ_e + H^2 Q_f \right) \right) \right|$$

JAI-CHAN HWANG AND HYERIM NOH PHYSICAL REVIEW D 71, 063536 (2005)







n_S





0.95

0.96

$n_{S} - 1 = 2(2\epsilon_{1} - \epsilon_{2} - \epsilon_{3}), \quad n_{T} = 2(\epsilon_{1} - \epsilon_{5}), \quad r = 16 \left| \frac{1}{Q_{t}} \left(\frac{c_{A}}{c_{T}} \right)^{3} \left(\epsilon_{1} - \frac{1}{4M_{p}^{2}H^{2}} \left(2Q_{c} + Q_{d} - HQ_{e} + H^{2}Q_{f} \right) \right) \right|.$ Planck TT, TE, EE + lowE+lensing 0.0045 0.0040 0.0036 0.9615 0.966 $N_{*} = 50$ $N_{*} = 60$ $\alpha = -1.4 \times 10^4$ $- \alpha = 0$ $- \alpha = 8 \times 10^3$ 0.97 0.99 0.98 1.00 n_S







GW170817 & GRB170817A: $-3 \times 10^{-15} \le c_T/c_{\gamma} - 1 \le 7 \times 10^{-16}$

Phys. Rev. Lett. 119, 161101 (2017) Astrophys. J. Lett. 848, L13 (2017)









Summary

- We studied Higgs inflation with a Gauss-Bonnet combination in the Einstein frame.
- to non-trivial effect(s).
- using the CMB data.
- The model is consistent with the CMB data for both positive and negative α .
- The propagation speed of GWs puts further constraints on the α .
- As a result, the positive values of the α is favored, i.e., $0 < \alpha \leq 3 \times 10^{-7}$.

• The Gauss-Bonnet coupling must be a function of a scalar field $\omega(\phi)$ to give rise

• Considering the relation $\omega(\phi) = \alpha \Omega^2(\phi)$, we put a constraint on the parameter α

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Thank you for your kind attention!

Backup Slides

The solutions to the background equations of motion:



(N)s 2 Field values at the end of inflation 55 60 65 70 50 $2. \times 10^{-15}$ $1. \times 10^{-15}$ 3-contribution $-1. \times 10^{-1}$ $-2. \times 10^{\circ}$ $-3. \times 10^{-1}$ $-4. \times 10^{-15}$ 55 60 65 70 75 80 50 55 60 65 70 75 80 50 55 60 65 70 $\ddot{s} + 3H\dot{s} + V_{,s} = -12c_1\xi_{,s}^{N}H^2(\dot{H} + H^2) + \frac{3}{2}\frac{c_2}{M_p^2}\left[H^2(\dot{\xi}\dot{s} + 2\xi\ddot{s}) + 2H\xi\dot{s}(2\dot{H} + 3H^2)\right] - \frac{1}{2}\frac{c_3}{M_p^3}\dot{s}\left[\ddot{\xi}\dot{s} + 3\dot{\xi}\ddot{s} - 6\xi(\dot{H}\dot{s} + 2H\ddot{s} + 3H^2\dot{s})\right].$



The solutions to the background equations of motion:



s(N) Field values at the end of inflation 55 50 60 65 70 N 10^{-13} 10^{-14} 10⁻¹⁵ 10^{-16} •••• C2 $\cdots C_3$ 10^{-17} $\ddot{s} + 3H\dot{s} + V_{,s} = -12c_1\xi_{,s}H^2(\dot{H} + H^2) + \frac{3}{2}\frac{c_2}{M_p^2}\left[H^2(\dot{\xi}\dot{s} + 2\xi\ddot{s}) + 2H\xi\dot{s}(2\dot{H} + 3H^2)\right] - \frac{1}{2}\frac{c_3}{M_p^3}\dot{s}\left[\ddot{\xi}\dot{s} + 3\dot{\xi}\ddot{s} - 6\xi(\dot{H}\dot{s} + 2H\ddot{s} + 3H^2\dot{s})\right].$



